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MASSACHUSETTS UNIV AMHERST ASTRONOMY RESEARCH FACILITY F/G 4/1
STUDY OF BACKGROUND RADIANCE IN UPPER ATMOSPHERE, (U) F19628-75-C-0187
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STUDY OF BACKGROUND RADIANCE IN UPPER ATMOSPHERE

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January 1980

Final Report

1 July 1975 - 30 September 1979

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This research was sponsored in part by the Defense Nuclear Agency
Subtask I25AAHIG32, Work Unit 42, entitled "SPINE Instrumentation."

AIR FORCE GEOPHYSICS LABORATORY
AIR FORCE SYSTEMS COMMAND
UNITED STATES AIR FORCE
WRIGHT-PATTERSON AIR FORCE BASE, OHIO 45433-6531

80-12-1-002

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SECURITY CLASSIFICATION OF THIS PAGE (When Data Entered)

19) REPORT DOCUMENTATION PAGE		READ INSTRUCTIONS BEFORE COMPLETING FORM
1. REPORT NUMBER <i>(18) AFGL-TR-80-0048</i>	2. GOVT ACCESSION NO. <i>AD-A092, 934</i>	3. RECIPIENT'S CATALOG NUMBER
4. TITLE (and Subtitle) <i>STUDY OF BACKGROUND RADIANCE IN UPPER ATMOSPHERE.</i>	5. TYPE OF REPORT & PERIOD COVERED <i>Final Rept 1 Jul 1975 to 30 Sep 79</i>	6. PERFORMING ORG. REPORT NUMBER <i>(14) UMASS-ARF-80-308</i>
7. AUTHOR(s) <i>(10) Hajime Sakai</i>	8. CONTRACTOR OR GRANT NUMBER(s) <i>(12) (15) F19628-75-C-0187</i>	9. PROGRAM ELEMENT, PROJECT, TASK AREA & WORK UNIT NUMBERS <i>PE61102F 2310G404 & 2310G408</i>
11. CONTROLLING OFFICE NAME AND ADDRESS <i>Air Force Geophysics Laboratory Hanscom AFB, Massachusetts 01731 Monitor/Richard Nadile/OPR-1</i>	12. REPORT DATE <i>(11) January 1980</i>	13. NUMBER OF PAGES <i>85</i>
14. MONITORING AGENCY NAME & ADDRESS (if different from Controlling Office) <i>(16) AFGL-TR-80-0048</i>	15. SECURITY CLASS. (of this report) <i>Unclassified</i>	16a. DECLASSIFICATION/DOWNGRADING SCHEDULE
16. DISTRIBUTION STATEMENT (of this Report) Approved for public release; distribution unlimited.		
17. DISTRIBUTION STATEMENT (of the abstract entered in Block 20, if different from Report)		
18. SUPPLEMENTARY NOTES This research was sponsored in part by the Defense Nuclear Agency Subtask I25AAXHX632, Work Unit 42, entitled "SPIRE Instrumentation."		
19. KEY WORDS (Continue on reverse side if necessary and identify by block number) Atmospheric emission Radiative transfer Scattering Infrared		
20. ABSTRACT (Continue on reverse side if necessary and identify by block number) The final report on the analysis of the SPIRE data. The major CO ₂ IR bands are analyzed by using a simplified model.		

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Introduction

The work of this contract was intended to develop the atmospheric model which integrates the emission and the scattering term in a combined form and to provide analysis of the earth atmospheric emission data gathered by the SPIRE payload in these two terms.

Our work performed throughout the entire period can be divided into five areas:

- (1) incorporation of various computer programs designed to calculate the atmospheric infrared emission level under a specific condition;
- (2) improvement of the algorithm employed in these programs;
- (3) development of an atmospheric infrared radiance model which unifies the emission and scattering term;
- (4) application and evaluation of the developed model to analysis of the SPIRE data; and
- (5) evaluation of some preliminary test data for the coming AFGL balloon-borne atmospheric emission measurement experiment.

The SPIRE experiment¹ was carried out during 1978, and the collected data were made available to us in late 1978. Our effort during the early phase of the contract period, i.e., prior to the SPIRE launching, was focused in the area of (1) and (2). Our Scientific Report No. 1 summarizes our progress made during the early phase.² The present report will describe our work covered in the entire period, with a special emphasis on the progress made since the last report.

Upon inspection of the SPIRE data, it became obvious to us that the existing computation programs are much more elaborate than required in the

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analysis. A quick check on the observed scattering radiance level revealed that the single-scattering theory with the Rayleigh particles produced a result adequate for the analysis. The Monte Carlo code developed by Blattner and his co-workers^{3,4} is found unnecessary for the analysis, because the computation based on the single scattering process produced a result which agrees, reasonably well, with the SPIRE observed data. The existing computation programs treat the scattering process entirely separately from the absorption-emission process. We find them rather inadequate for explaining why some molecular bands change their absorptive behavior against the scattering background to become emissive as the altitude increases. A theory which unifies both processes in the radiative formulation is definitely needed for the analysis. A limit placed on time restricted us to develop a computation algorithm which generates the spectral radiance level as a function of frequency and an observation condition in an automatic fashion. We feel that we are on the right track for seeking a solution to the problem. Nonetheless, we came to understand the basic radiative mechanism of the upper atmosphere with an extent that we can place a proper perspective to the analysis of the SPIRE data.

Radiative Process

The radiative transfer in the atmosphere is traditionally analyzed in treating the infrared radiation field as a continuous flow of radiative energy. The interactions between the radiation field and the atmospheric gas is formulated in an integral equation. Several computer programs have been written for analysis of the radiation transfer problem. Our approach taken in the present study is slightly different from the traditional analysis. We tried to analyze the radiative transfer in the upper atmosphere in following the interactive process between the photons and molecules.

The photons incident into a layer of atmosphere would go through one of three processes: (1) they may be absorbed by the atmospheric molecules or atoms; (2) scattered by the Rayleigh particles or by the aerosols; and (3) left unperturbed. We can formulate these interactions between the photon field and the molecular system in terms of the density of these two interacting partners and a quantity which indicates a strength of the interaction.

The interaction between the photons and the molecules (or the atoms) in the absorption process is specified by Einstein's formulation.⁸ The number of molecular transitions N_t in unit time per unit volume, when interacting with the photons of a frequency ν , is given by

$$N_t = (N_\ell B_{\ell u} - N_u B_{u \ell}) \rho(\nu). \quad (1)$$

In this expression, the following notations are used:

N_ℓ : number of molecules per unit volume in lower state ℓ

N_u : number of molecules per unit volume in upper state u

$B_{\ell u}$: Einstein coefficient for induced absorption

$B_{u \ell}$: Einstein coefficient for induced emission

$\rho(\nu)$: spectral density of the radiation field in unit volume at frequency ν

The density $\rho(\nu)$ is related to the photon density N_p given in unit volume per second by

$$\frac{c}{4} \rho(\nu) = N_p h\nu, \quad (2)$$

where h is the Planck constant and c is the speed of light. Combining Eqs. (1) and (2), the molecular transitions N_t is given by

$$N_t = (N_\ell B_{\ell u} - N_u B_{u \ell}) \frac{4N_p h\nu}{c}. \quad (3)$$

This expression can be rewritten by using the molecular absorption

strength $S_{\ell u}$ (cross-section)⁹

$$N_t = 4N_p N_o S_{\ell u}, \quad (4)$$

where

$$S_{\ell u} = \frac{h\nu(N_{\ell}^B - N_u^B)_{\ell u}}{N_o c} \approx \frac{h\nu}{c} B_{\ell u} \left(1 - \frac{N_u}{N_{\ell}}\right) \quad (5)$$

with N_o the number of molecules in unit volume. In deriving Eq. (5), it is assumed that

$$N_o \approx N_{\ell}. \quad (6)$$

The Eq. (4) derived above specifies the number of molecular transitions N_t in the photon density N_p , the molecular density N_o , and the molecular absorption strength $S_{\ell u}$ which is experimentally determined in the laboratory.

The scattering process can be formulated in the same way.⁴ A traditional approach is somewhat different in that the incident radiation field is given in the irradiance; the power scattered toward the direction ϕ by a molecule is given by

$$\sigma_m(\phi) E_o = I(\phi), \quad (7)$$

where E_o is the irradiance of the incident plane wave. We can formulate the incident irradiance by

$$E_o = N_p h\nu, \quad (8)$$

again using N_p , the photon density in unit volume per unit second. Now we find that the power scattered by a molecule is given by

$$I(\phi) = N_p h\nu \sigma_m(\phi). \quad (9)$$

The radiative power scattered toward the direction ϕ per unit solid angle by an ensemble of molecules N_o in unit volume is given by

$$I'(\phi) = N_o h\nu N_p \sigma_m(\phi). \quad (10)$$

Thus we get the desired form for the number of photons scattered toward the

direction ϕ ;

$$N_p(\phi) = N_o N_p \sigma_m(\phi). \quad (11)$$

For the Rayleigh scattering, the scattering cross-section is given by

$$\sigma_m(\phi) = \frac{9\pi^2 v^4}{N^2 c^4} \left(\frac{n^2 - 1}{n^2 + 2} \right)^2 \sin^2 \phi \quad (12)$$

$$\approx \frac{4\pi^2 v^2}{N^2 c^2} (n-1)^2 \sin^2 \phi$$

where N is the number of molecules in a unit volume at S.T.P.⁴

The energy absorbed by a molecule from the radiative field is sooner or later released from it in a de-excitation process. In the upper atmosphere, the de-excitation is done by the radiative process. The photon absorbed by a molecule is re-emitted to the radiation field isotropically. If the original absorption occurs between the ground state and the first excited vibrational state, there is no significant change in the photon energy of the absorbed and of the re-emitted. In the lower atmosphere, a local thermal equilibrium is established; the energy absorbed by a molecule is dissipated into the kinetics energy shared in the molecular ensemble. The interaction between the molecular ensemble and the radiation field under the local thermal equilibrium condition is treated below in calculation of the photon density. The de-excitation under this condition is done predominantly by the collisional process. In a transient region from the lower to the upper atmosphere, the molecular system de-excites by both processes.

Photon Density in Atmosphere

There are two radiation sources which must be considered in the infrared

radiative process of the atmosphere: the solar radiation penetrating the atmosphere from the top, and the radiation going upward to escape to the outerspace from the earth's surface. The energy density $\rho(v)dv$ of the black-body at temperature T is given by

$$\rho(v)dv = \frac{8\pi h v^3}{c^3} \frac{dv}{e^{\frac{hv}{kT}} - 1}. \quad (13)$$

The photon density N_p in unit volume per unit second per wavenumber interval $d\sigma$ emitted by the same blackbody radiator is

$$N_p d\sigma = \frac{c^2}{4\pi h v} \rho(v)dv \\ = \frac{2\pi v^2}{c} \frac{d\sigma}{e^{\frac{hv}{kT}} - 1} = \frac{2\pi \sigma^2 c}{h c \sigma} \frac{d\sigma}{e^{\frac{hv}{kT}} - 1}. \quad (14)$$

For the telluric radiation, the photon density is calculated by taking T at a value between 225°K and 300°K. For the solar radiation, it is calculated by the same expression with an attenuation factor;

$$N_{p_s} = \left(\frac{r_s}{R_s}\right)^2 \frac{2\pi \sigma^2 c}{h c \sigma} \frac{d\sigma}{e^{\frac{hv}{kT}} - 1}, \quad (15)$$

where r_s is the solar radius and R_s is the solar distance, T is about 6000°K.

The front factor $(r_s/R_s)^2$ is approximately 2.2×10^{-5} . In Figure 1, the photon density for three cases is plotted as a function of the wavenumber σ : (1) the solar radiation calculated by Eq. (15), (2) a blackbody at $T = 300^\circ K$, and (3) a blackbody at $T = 225^\circ K$. The telluric photon density in the upper atmosphere is affected by the optical thickness of the lower atmosphere. In some spectral ranges, there is no atmospheric molecular absorption, and the photon density for the $300^\circ K$ blackbody remains unaffected from the earth's surface to the top of atmosphere. In a region where the atmospheric

absorption is extreme, the photon density in the upper atmosphere is controlled by the energy density of the blackbody radiation at the tropopause temperature $T = 225^{\circ}\text{K}$. We can observe in Figure 1 that the molecules and the atoms in the upper atmosphere are excited by two types of radiations which have a vastly different spectral characteristic. The telluric photon density remains unaffected by the solar daily motion. The photons in a region below 2000 cm^{-1} are predominantly telluric, while those in a region above 2000 cm^{-1} in the daytime are solar.

The emission due to the CO_2 transitions, one for the $(01101-00001)$ ¹⁰ band at 670 cm^{-1} (the v_2 band) and another for the $(00011-00001)$ band at 2300 cm^{-1} (the v_3 band), is ready for a qualitative analysis using the photon density calculated for the curves shown in Figure 1. (A more detailed quantitative analysis will follow later in the report.) Once accepting the photon spectrum shown in Figure 1, we find that the photons available for exciting the v_2 band are telluric. The temporal characteristics of the telluric photons remain independent of the daily cycle. Thus the v_2 ($01101-00001$) band emission at 670 cm^{-1} must remain unchanged between the day and the night time. The SPIRE data shown in Figure 2 for this CO_2 band confirms a stationary nature of the emission. The $\text{CO}_2 v_3$ ($00011-00001$) band at day side is excited by the solar photons, and by the telluric photons at night side. Since the v_3 transition is very strong, the telluric photons available in the upper atmosphere are those of a 225°K blackbody. The difference in these two densities between the solar and the telluric is two orders of magnitude. Thus the radiance level of the transition should undergo a change of about 2 orders of magnitude between the day and the night. The SPIRE data in Figure 3 indicates that the v_3 band emission primarily undergoes the diurnal change described above. Upon a careful inspection of the data, the radiance level at nighttime is brighter than what can be predicted by the telluric photon level.¹⁷ We will make an analysis on this anomaly later in the report.

Photon Density / $\text{cm}^3 \text{sec} \text{cm}^{-1}$

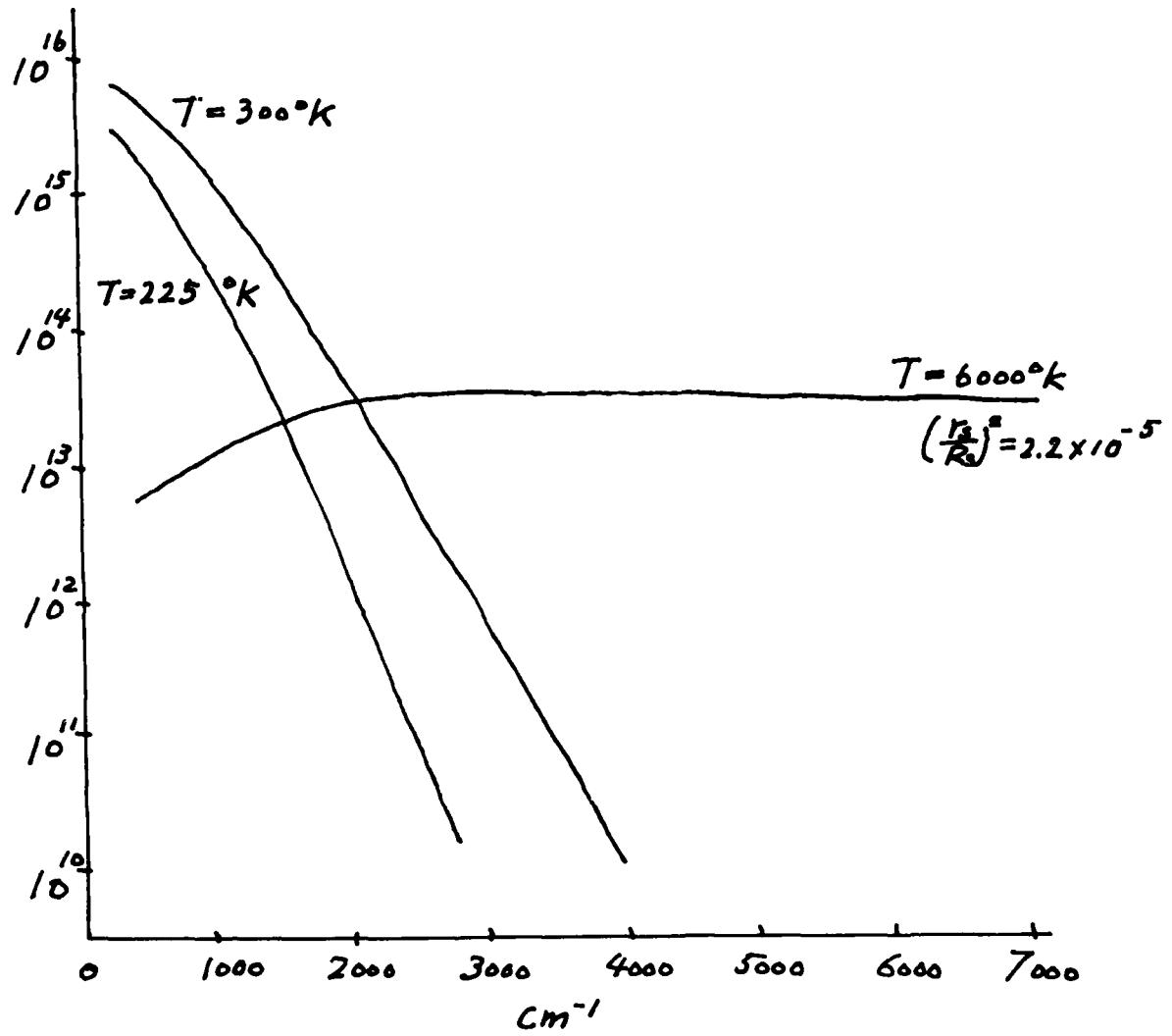


Figure 1

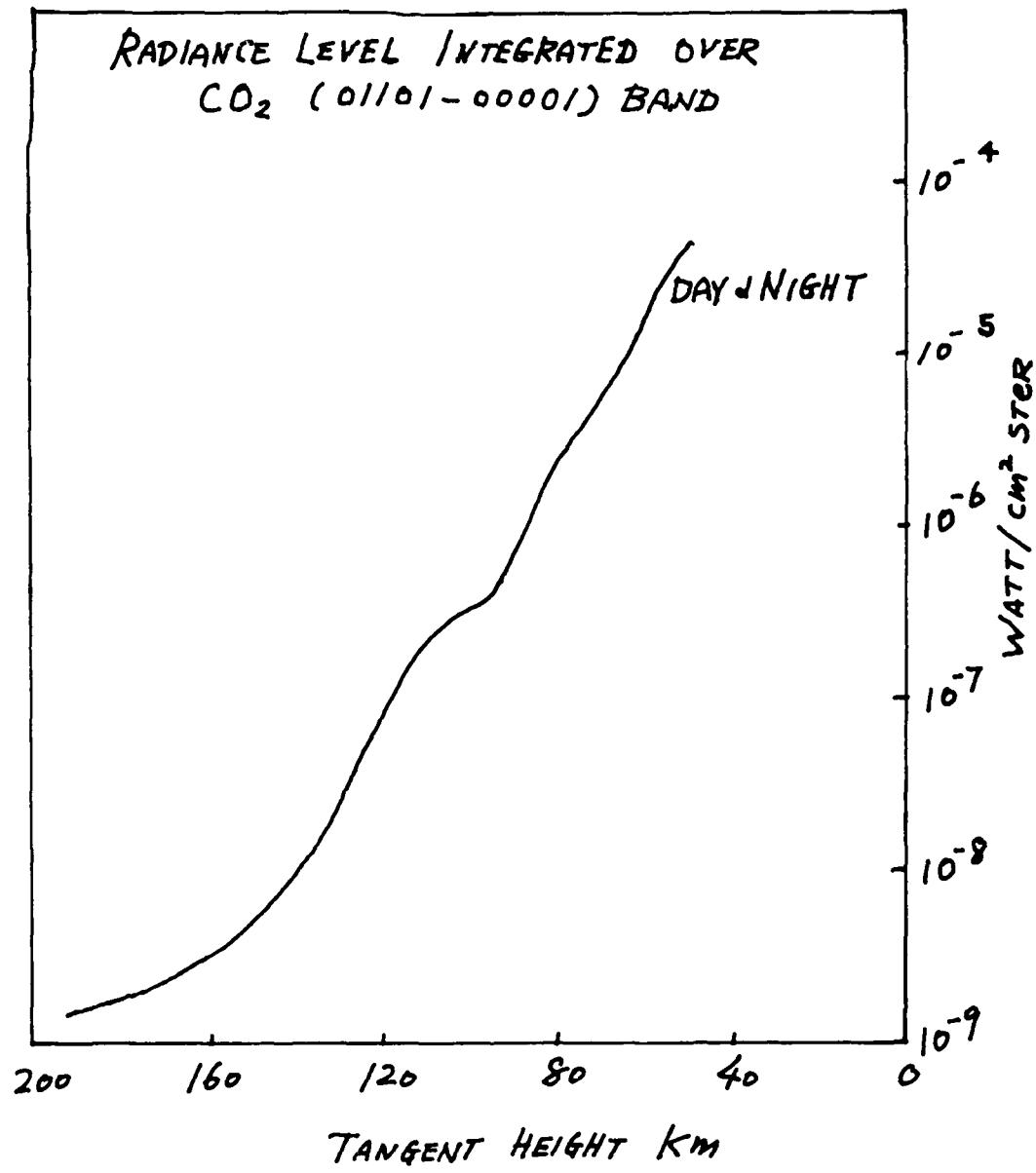


Figure 2

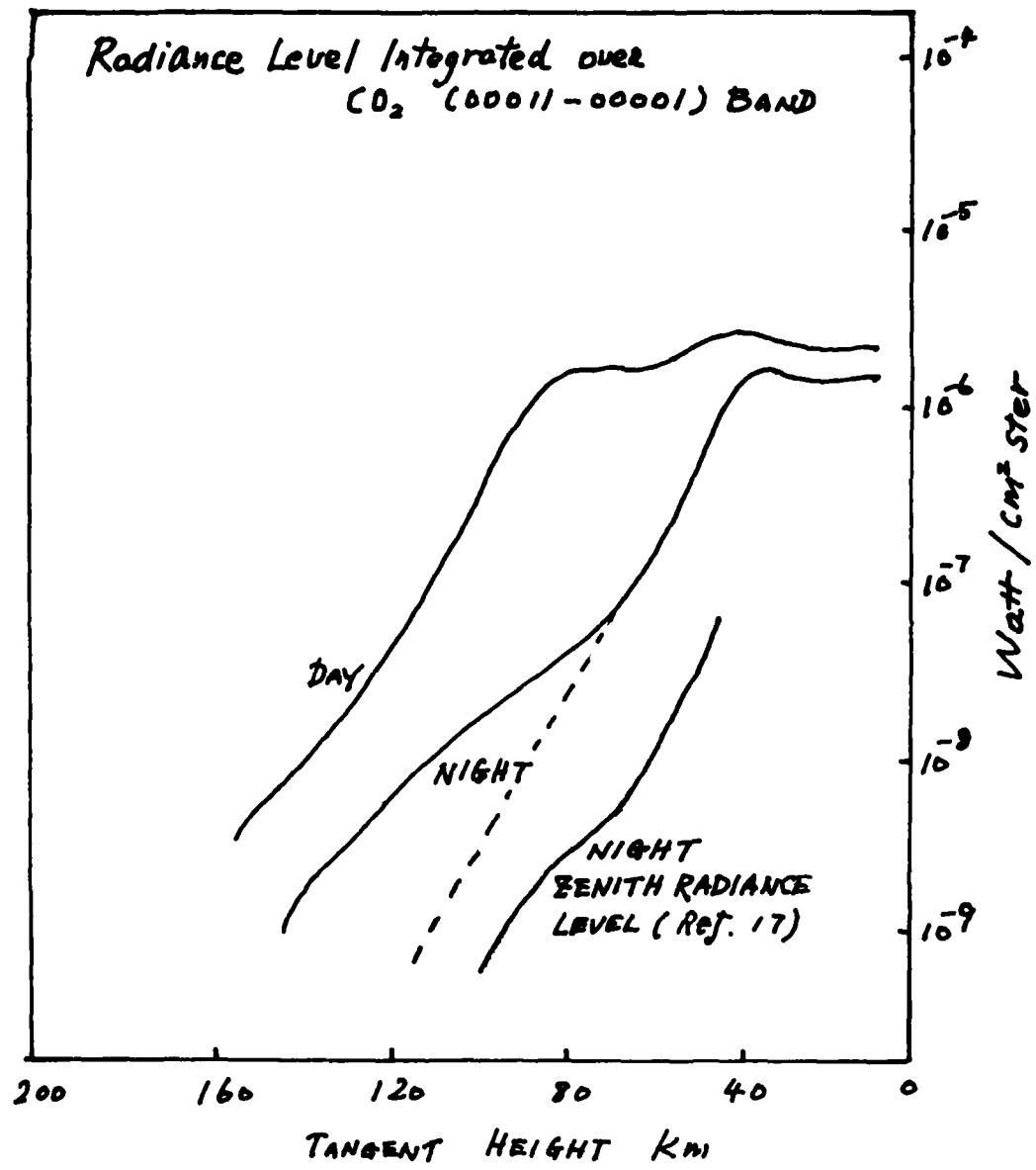


Figure 3

Model for Calculating Atmospheric Radiance Level

The expressions obtained above in Eqs. (4) and (11) are concerned with an individual interacting process in which a single photon is either absorbed or scattered by a molecule. In this section we will develop a model which relates the atmospheric radiance level to the basic interaction process given by these expressions.

The energy absorbed by a molecule in interaction with the photon field must be released in a de-excitation process. In the upper atmosphere, the only process available for de-excitation is a radiative process. A question which must be raised and answered for such a radiative excitation-de-excitation process concerns an establishment of equilibrium population density among various molecular states. Degges and other workers developed the computation program capable of giving an estimate of the radiative level in the normal atmospheric environment at high altitude.^{11,12} We are truly thankful for the existence of these programs which provided a valuable help to develop our insight into the problem.

We modified the Degges program to suit our needs, and to improve the computation efficiency in our CYBER system. A brief description of our effort along this direction will be given later in this report, together with the listing of our modifications. The atmospheric model is judged on its ability for producing a reasonable estimate of the infrared radiance level; in particular, it is questioned on precision and practicality.

The computation scheme to obtain the radiance level must find the equilibrium population distribution which is established among the vibrational states when the molecular system is subjected to balance with the photon field. The precision needed in the computation is somewhat controlled by a spectral resolution of the result. The emission involving

the hot-band transition can be distinguished only with an improved spectral resolution, because it is usually buried under the fundamental transition. We decided to ignore these hot-band transitions from our model in a belief their contribution is not significant to the computation. The assumptions made in our model* are certainly open for debate and all criticisms to our viewpoint are welcome. Our view that all transitions observed in the data occurred from the ground state, in turn, simplified the calculation to a great extent. A direct correlation was resulted in our formulation between the radiance level for a particular tangent height and the molecular density of the concerned molecular system at this altitude.

The total energy emitted by the molecules in a unit volume through the radiative de-excitation is equal to the energy absorbed by the same molecules if that is the only process feasible. The absorbed energy is given by the number of transition times the photon energy: $b' = h\nu N_t$. The energy released in the radiative de-excitation is emitted isotropically. Thus the energy re-emitted per a unit area over a unit solid angle is given by: $b = b'/4\pi = h\nu N_t/4\pi$. The photons released from a molecule are very unlikely absorbed again, since the atmosphere is optically thin.[†] Even if they are absorbed, they are re-emitted in the photon field. The photons which reach a sensor of a unit area with a field view of a unit solid angle are computed simply by accumulating the contribution from each unit volume which falls into a line of sight. Since the optical étendue preserves, the summation which we must take in the calculation is equal to that of each contribution b taken over a column of a unit cross-section extending from the sensor to the space all the way along the line of sight. The radiance level per unit solid angle integrated over the band is then given by

*See the Appendix D.

[†]The absorption along the limb-view path of 100 km tangent height is less than 5% for the (00011-00001) band of CO₂.

$$B = \frac{h\nu N_t}{4\pi} = \frac{h\nu S_u l N N_p}{\pi} \quad (16)$$

where N is the total number of molecules counted along the line-of-sight per unit area;

$$N = \int N_o dl . \quad (17)$$

The density of atmospheric molecules falls off approximately exponentially with the altitude. If we can assume the density $N(h)$ as a function of altitude h given by

$$N(h) = N(h_o) e^{-\frac{h-h_o}{H}} \quad (18)$$

with a characteristic constant H , the total number of molecules between two altitudes h and h_o is given by

$$\begin{aligned} N(h, h_o) &= \int_{h_o}^h N(h) dh = N(h_o) \int_{h_o}^h e^{-\frac{h-h_o}{H}} dh \\ &\approx N(h_o) H \end{aligned} \quad (19)$$

For the limb view calculation, we need to calculate a distance l between two altitudes, the tangent height x and a height above by H . The distance l is calculated by

$$l = (x+H+\rho) \sqrt{1 - \left(\frac{x+H}{x+H+\rho}\right)^2} \approx \sqrt{(2H)(x+\rho)} \quad (20)$$

where ρ is the earth radius (≈ 6360 km). To calculate the total density N along a line of sight to the tangent height x , we use the approximation Eqs. (19) and (20):

$$N = N_o(x) 2 \sqrt{2(H)(x+\rho)} \quad (21)$$

where $N_o(x)$ is the molecular density at altitude x . With the approximate total density N derived in Eq. (21) the limb-view radiance at the tangent

height x , is given by an expression

$$B = \frac{hv}{\pi} S_{ul} N_o(x) 2 \sqrt{2H(x+\rho)} N_p. \quad (22)$$

The expression given by the equation above requires an estimate of the photon density N_p . The value shown in Figure 1 is given in a unit wavenumber interval. The photons which contribute to the transition are those within the width of the vibrational-rotational lines. If we take the photon density to be the value over the band, say 100 cm^{-1} for the CO_2 (01101-00001) transition, we obviously get an erroneous result. The effective photon density must be adjusted by multiplying an effective number of the rotational lines within the band times the line width, which we can safely assume the Doppler width for those in the upper atmosphere.

CO_2 Bands, (01101-00001) and (00011-00001)

The CO_2 molecule is relatively stable in the atmosphere, because neither electronic transitions nor photo-dissociation occurs in the entire spectral range below the Schuman-Runge Continuum of O_2 . The mixing ratio remains constant in the atmosphere below 120 km where the O_2 absorption is very effective in blocking those high energy photons in the extreme uv region.

As a result, the equilibrium population distribution is established among various vibrational levels in the electronic ground state with insignificant perturbation from higher excited electronic states. The equilibrium is achieved in balance with the infrared photon field. Under this condition a major concern must be directed to the equilibrium population established for the metastable (10001) and (10002) level, which locates approximately 1300 cm^{-1} above the vibrational ground level. Since there are no optical transitions allowed between those and the ground level, a primary

candidate for pumping these levels is the radiationless collisional process which occurs only in the lower atmosphere. The SPIRE data indicate that the emission by the (10001-01101) and (10002-01101) transitions is definitely observable at low tangent height in both the day and night conditions. At high tangent height both of the emissions become less distinctive. It may be concluded that the radiative pumping mechanism for these metastable states is rather minor. The observation coincides with the results obtained by the Degges' radiation balance computation program.

The limb-view radiative levels of the observable CO₂ bands for high-tangent height are calculated with a reasonable accuracy by the model developed above. There are two clearly observable CO₂ emissions in the SPIRE data, one for the (01101-00001) band and another for the (00011-00001). As mentioned above, the emissions by the (10001-01101) and the (10002-01101) transitions are observable in the low-tangent-height data. Both the (10011-00001) and the (10012-00001) band will be discussed together with the scattering continuum observable in the 3000 cm⁻¹ ~ 10,000 cm⁻¹ region.

The radiance levels shown in Figures 2 and 3 are extrapolated from the SPIRE data by integrating them over the entire band. At a tangent height of 100 km, the (01101-00001) emission is 3×10^{-7} (watt/cm² sterad). Using the expression given in Eq. (22):

$$B = \frac{hv S N_o N_p}{\pi}$$

our calculation yields 2.0×10^{-7} watt/cm² sterad. In deriving the value, we take

$$hv = 1.32 \times 10^{-20} \text{ Joule}$$

$$S = 8.3 \times 10^{-8} \text{ mol}^{-1}/\text{cm}^2 \text{ cm}^{-1}$$

$$N_o = 5 \times 10^7 \times 10^9 \text{ for the } 100 \text{ km tangent height, and}$$

$$N_p = 3 \times 10^{15} \times \Delta\sigma \times n = 3 \times 10^{15} \times .001 \times 40,$$

the Doppler width $\Delta\sigma$ calculated by

$$\Delta\sigma = \sigma_0 \frac{\sqrt{3kT}}{mc^2} \approx 0.001 \text{ cm}^{-1} \text{ for } \sigma_0 = 670 \text{ cm}^{-1},$$

and 40 rotational lines as those effectively emitted.

The SPIRE data indicate that the (00011-00001) emission at 100 km tangent height is 3×10^{-7} (watt/cm² sterad) for the daytime level and 3×10^{-8} (watt/cm² sterad) for the nighttime level. Our simple-minded calculation shows that $B = 2.1 \times 10^{-7}$ (watt/cm² sterad) for the day data and 2.1×10^{-9} (watt/cm² sterad) for the night data. In deriving these figures, we take

$$hv = 4 \times 10^{-20} \text{ Joule}$$

$$S = 9.6 \times 10^{-17} \text{ mol}^{-1} / \text{cm}^2 \text{ cm}^{-1}$$

$$N_o = 5 \times 10^7 \times 10^9$$

$$N_p = 3 \times 10^{13} \times .1 \text{ for the day and } 3 \times 10^{11} \times .1 \text{ for the night.}$$

The observed and the calculated data for the day level agree reasonably well.

The observed nighttime radiance level is about one order of magnitude higher than our calculation. In Figure 3, a dotted curve shows the nighttime radiance level calculated by our formulation, Eq. (22). Kumer et al¹⁷ analyzed the nighttime zenith radiance data of the same band (00011-00001), which is shown in the same Figure. Both nighttime calculations are remarkably similar.

The SPIRE data differ from the data used in the Kumer's analysis in that they are for the limb view. The total CO₂ molecules involved in both data differ by an order of a magnitude. Thus, the radiance level differs by the same magnitude. The deviation from the exponential decay line in the radiance level occurs approximately at the same altitude for both data. Kumer et al interpreted this anomaly in terms of the extra pumping produced by the vibrationally excited N₂ molecules. The SPIRE data indicates that the

difference of the nighttime radiance level from the earthshine figure is more pronounced in the high altitude region where the interaction to the neighbor molecules should decrease. The anomaly in the nighttime radiance level for this band remains mysterious. We studied a possibility of observing the emission by the $W^3\Delta u - B^3\Pi g$ transition of N_2 which could be more intense at the altitude above 100 km because of a higher electron density there. A radiance level for the transition estimated on a basis of the known data fails to support the possibility.

Since at a low tangent height the band is completely saturated and the line structure is lost, the photon density available for the absorption is no longer adjusted by the effective number of the rotational lines excited in the band times their width. It is given by the width of the entire band times the photon density: $200 \times 3 \times 10^{11}$ photons/cm²/sec. Taking the photon energy $h\nu = 4.6 \times 10^{-20}$ Joule, we can estimate that the saturation level is given by

$$\frac{N_p h\nu}{\pi} = 0.9 \times 10^{-6} \text{ (watt/cm}^2 \text{ sterad).}$$

Scattering Radiance Observable Between 6000 cm^{-1} and 7500 cm^{-1} and the
 $\text{CO}_2\text{-H}_2\text{O}$ Band at 3600 cm^{-1}

The radiance level in a spectral region above 4000 cm^{-1} is strongly influenced by the scattering process, in particular the level observable at a low tangent height. For the Rayleigh process the scattering differential cross-section $\sigma(\phi)$ is given by Eq. (12):

$$\sigma_m(\phi) = \frac{9\pi^2\sigma^4}{N^2} \left(\frac{n^2-1}{n^2+1} \right)^2 \sin^2 \phi \approx 4\pi^2\sigma^4 \left(\frac{n-1}{N} \right)^2 \sin^2 \phi,$$

where σ is the optical frequency in cm^{-1} , both n the refractive index of the atmosphere and N the number density of the air molecule are those values at the S.T.P. condition, and ϕ is the scattering angle. The factor $(\frac{n-1}{N})^2$ is a slow varying function of σ . The scattering cross-section is thus predominantly controlled by σ and ϕ . Figure 4 shows the front factor $4\pi\sigma^2(n-1/N)^2$ as a function of σ . The radiance level obtained by the SPIRE experiment is compared with the computed value for the Rayleigh scattering in Figure 5. The observed values are replotted as a function of σ and of a tangent height in a unit of $\text{watt}/\text{cm}^2 \text{ strad cm}^{-1}$. Irregular structures which can be considered as the molecular absorption or emission are removed by smoothing out. In the computed values, N_p is assumed by the solar value, $3 \times 10^{13} \text{ photons}/\text{cm}^3/\text{S}/\text{cm}^{-1}$. The observed radiative level supports the computed molecular scattering level even though a minor difference exists, dependent on the tangent height and the spectral frequency.

The comparison shown in Figure 5 leads us to conclude two characteristic features of the radiance level in the $3000 \sim 7500\text{ cm}^{-1}$ range: (1) the scattering radiance is controlled by the single-scattering process; and (2) its main mechanism is the Rayleigh process. The aerosol particles contribute to the over-all scattering process, not significantly even at a low tangent height of 15 km.

A problem which we will discuss below plays a central importance in calculation of the infrared radiance level in the upper atmosphere. It is a combined effect of the molecular absorption-emission process and the scattering process. It was treated rather lightly, if not completely side-stepped. The computation modeling of the atmospheric radiance developed to cover a region where only one of these processes is important in the radiative process. There were no models developed to treat the situation where the processes are competitive. The data obtained by the SPIRE experiment is in fact the first to observe the radiance level generated by these processes in competition. The data revealed an extremely remarkable feature for the radiance level in the region of the $\text{H}_2\text{O}-\text{CO}_2$ 2.7μ band ($3600 \sim 3800 \text{ cm}^{-1}$). The radiative level observed in the 15 km tangent height data shows an absorptive feature for the band, while that in other data shows an emissive feature. The radiance level remains rather stationary at 10^{-7} (watt/cm² strad μ) or 7.5×10^{-11} (watt/cm² strad cm^{-1}) over a wide range of the tangent height from 15 km to 41 km.

By means of the molecular absorption process the molecule is excited to a higher vibrational state in interacting with the radiation field. The energy absorbed by the molecule in the excitation to a higher vibrational state is equal to that supplied by the photon field. During the de-excitation stage, the same amount of energy is released. Dependent on the environmental surround to the molecule, the energy release takes place to the photon field or to others. If the interaction with the photon field dominates the de-excitation process, there is no net energy loss in the photon ensemble. The absorbed energy from the photon field is released back to the photon field by the de-excitation process. When the collisional process dominates the de-excitation process, the energy absorbed from the photon ensemble is

not released back. The photon density in the ensemble decreases until the photon ensemble reaches a thermal equilibrium with the molecular ensemble. The photon density which is established under a thermal equilibrium is given by the expression which was used to compute the values for Figure 1:

$$N_p = \frac{2\pi\sigma^2 c}{\frac{h\nu c}{kT}} e^{-\frac{h\nu}{kT}}$$

The radiative level observable from such an ensemble of the photons and the molecules is given by

$$I = \frac{h\nu N_p}{\pi}.$$

The stationary emission observed for the 2.7μ H_2O-CO_2 band corresponds to the situation described above. What we can observe is the photons generated by the H_2O-CO_2 band which are in a thermal equilibrium at the tropopause temperature. Using the blackbody photon density $N_p = .92 \times 10^7$ at 3600 cm^{-1} and $225^\circ K$, and the photon energy $h\nu = 7.1 \times 10^{-20}$ Joule at 3600 cm^{-1} , the saturation level I is calculated as 2.1×10^{-11} watt/ cm^2 sterad cm^{-1} . With assuming $T = 260^\circ K$, the computation yields 5×10^{-11} watt/ cm^2 sterad cm^{-1} . This value is independently determined from the internal excitation-de-excitation mechanism of the molecular system. Whenever the molecular system establishes a thermal equilibrium with the environment under the saturation absorption condition, the photon field must be in an equilibrium with it. The stationary photon flux observed for the 2.7μ H_2O-CO_2 band represents this situation, as was the case for the CO_2 (00011-00001) transition at lower altitude. If the saturation photon flux is larger than the scattering photon flux, the radiance level of the band in question is emissive. If it is smaller, the radiance level is absorptive.

$$4\pi^2 \left(\frac{n-1}{N}\right)^2 \alpha^4$$

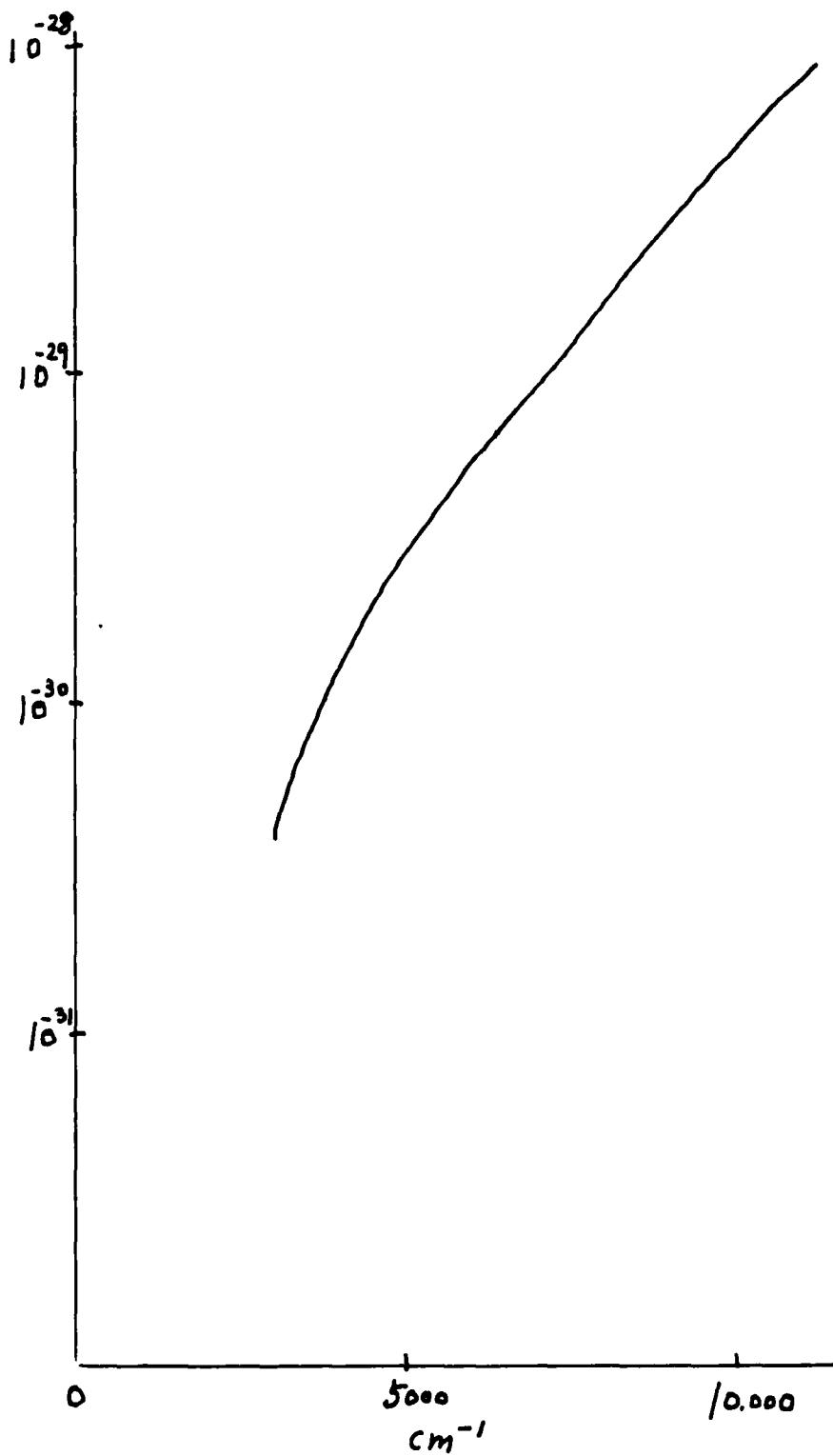


Figure 4

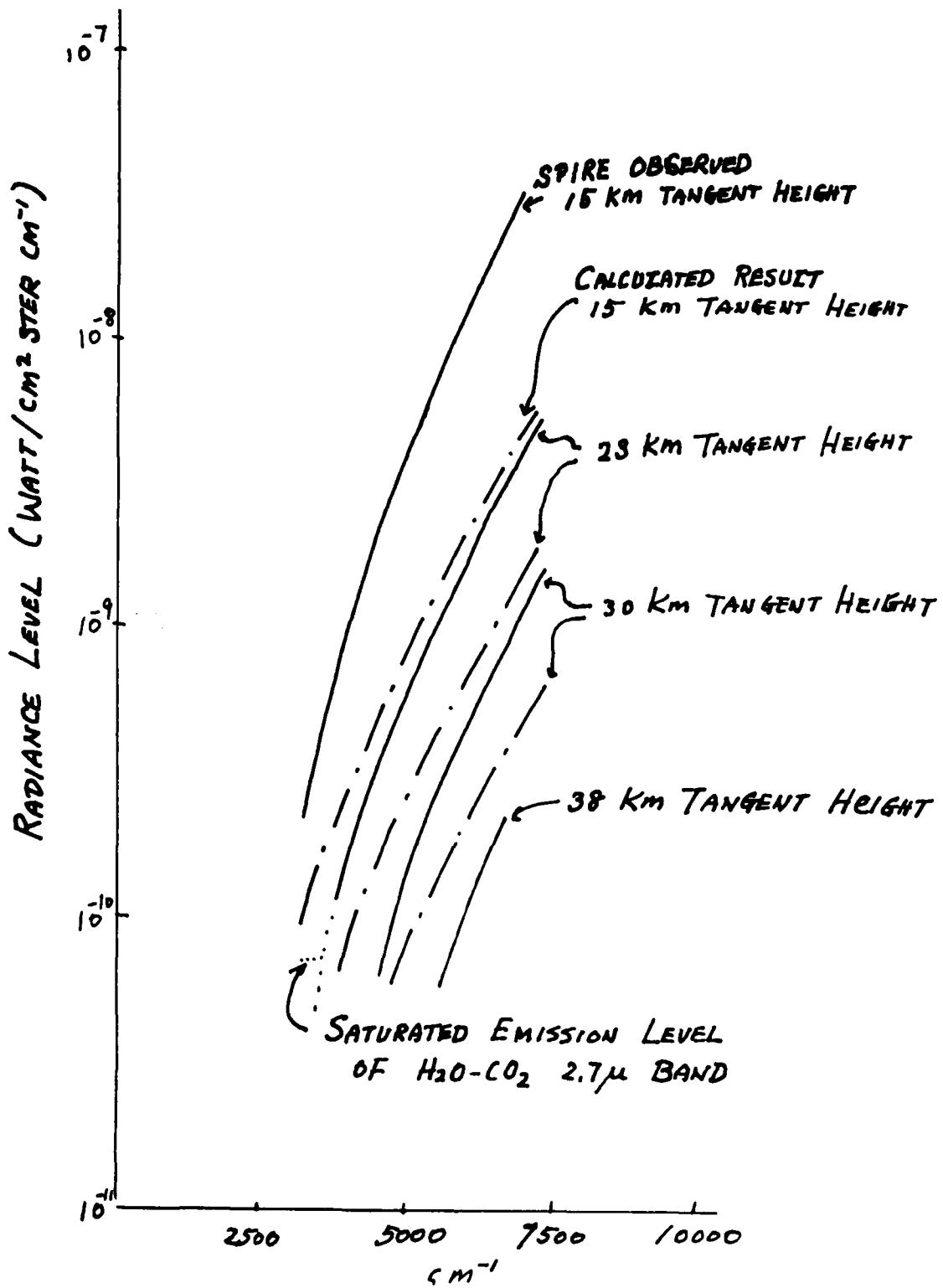


Figure 5

In both cases, the radiation level remains constant.

The Radiance Level of a Molecular Band

In the last two chapters, we discussed the saturated radiance level of the v_3 band of CO_2 and of the $2.7\mu \text{ H}_2\text{O}-\text{CO}_2$ band. We found that the photon flux under the saturation conditions is given by the blackbody radiation formula; i.e., that the molecular system and the photon ensemble come to establish a thermal equilibrium. For a given molecular band, there exists a critical altitude that the photon density is completely specified by the blackbody radiation formula independent from the vibrational-rotational transition. In the atmosphere above the critical altitude, the photon density is specified by the molecular transition equation, Eq. (4):

$$N_t = 4 N_o S.$$

A sensor placed very high in the atmosphere collects the photons generated by the molecular transitions which occur along a line-of-sight. With the approximation of a thin optical density air, the total molecular transitions which contribute to the observation are given by integrating the transitions N_t along the path:

$$N = \int N_t d\ell . \quad (23)$$

Under the approximate expression derived for an observation at a tangent height x , we calculate the total density per a solid angle per a unit cross-section by

$$N(x) = \frac{4N_o S}{\pi} N_o(x) 2 \sqrt{2H(x+\rho)} . \quad (24)$$

In deriving the equation above, an assumption is made that the photon density N_p responsible for exciting the molecular transition is constant along the path. In a spectral region below 2000 cm^{-1} where the telluric photons at the tropopause temperature control the excitation, the assumption is obviously

well justified. In a higher cm^{-1} region where the solar photons are dominant, a blackbody temperature corresponding to the solar photon density is much higher than the tropopause temperature as seen in Figure 6. The radiance level of the molecular emission reaches a maximum plateau at the saturation photon density specified by the temperature of tropopause. Thus the molecular transitions, even when they are integrated along a long path, are much less than the solar photon density flux which has a temperature much higher than the saturation photon temperature. The loss caused by the molecular transition along the path is therefore much smaller than the total solar photon density, and a constant photon density is maintained, as assumed above. This justification obviously breaks down as the radiance level approaches the saturation level.

In a non-saturating region, we find that the observable photons given by Eq. (22) are controlled by the factor $N(x_0)$, the molecular density at the tangent height X . In the first-order approximation, the radiance level at a particular tangent height gives the density of the observing molecule at that altitude. For a stable molecule, its density decreases exponentially as a function of the altitude. We can model the radiance of a molecular band, as shown in Figure 7, by an exponential function with a plateau at the critical altitude. The model certainly agrees with the radiance data collected by the SPIRE for the v_3 band of CO_2 .* The computed results by the Degges program as well as by the LOWTRAN for various bands support this model.

*The atmospheric sounding technique currently experimented uses this model. It is to observe the saturation photon density at the critical altitude, which varies from a spectral region to another as a function of the absorption strength. The observed radiance level gives a local equilibrium temperature at the altitude. An atmospheric sounding program proposed to NESS will use the (01101-00001) and (00011-00001) transitions of CO_2 .¹³

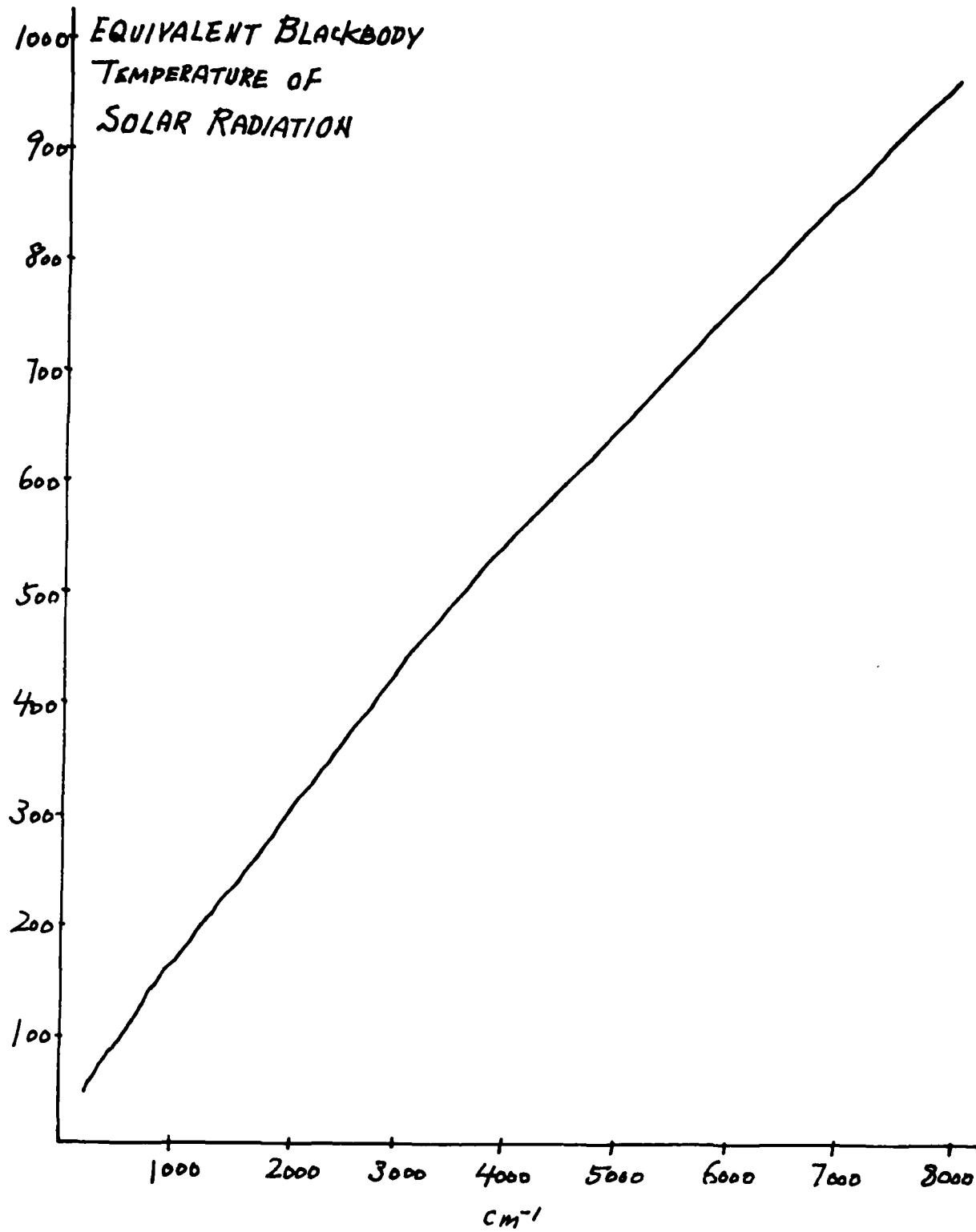


Figure 6

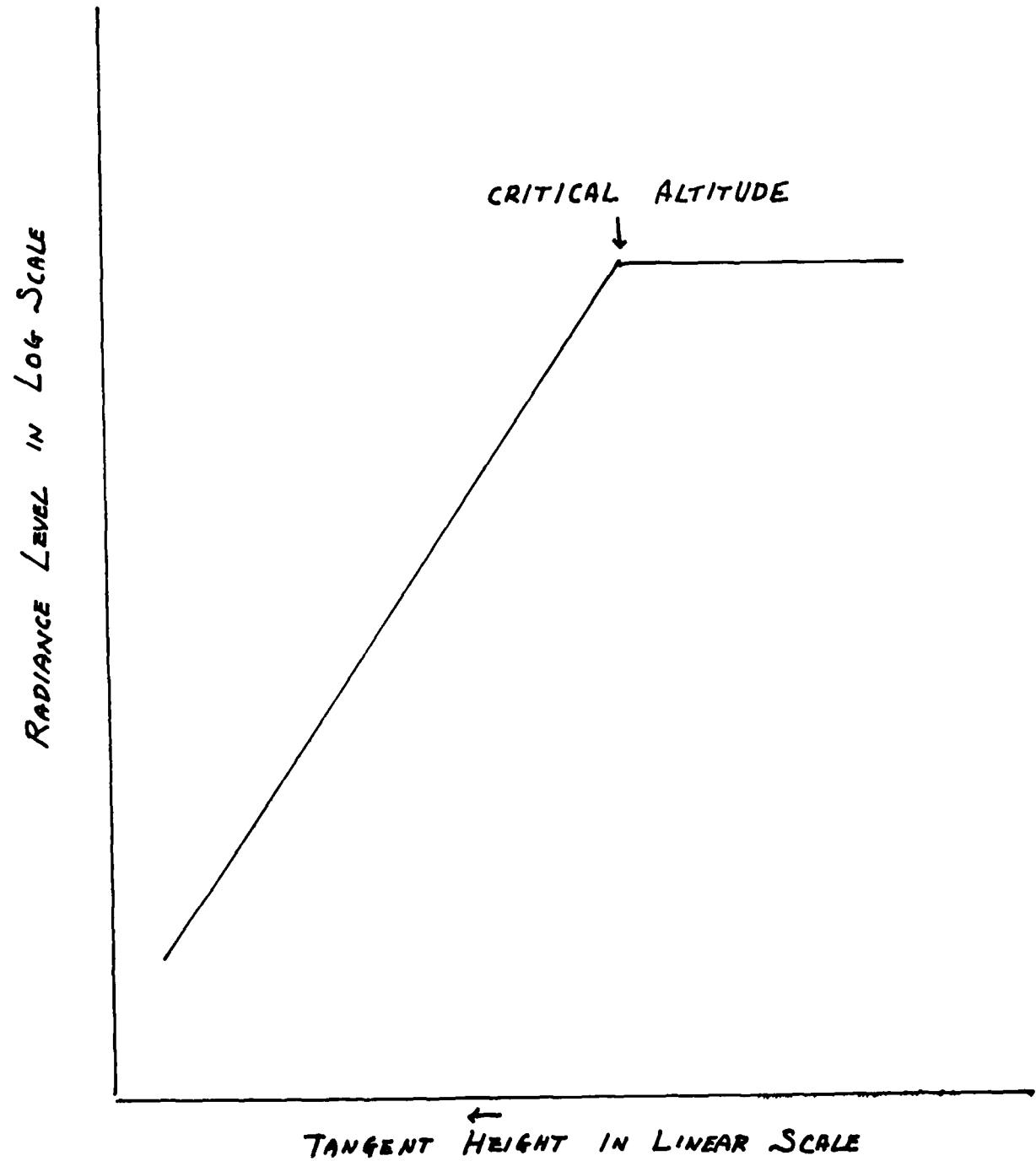


Figure 7

In the upper atmosphere, the radiance level is thus dictated directly by the molecular density. A deviation of the radiance level from the exponential decay is expected if the molecular concentration falls off rather differently from the exponential model. The SPIRE data on the infrared radiance level provide a base to determine the concentration of the observed atmospheric species as a function of the altitude. From our point of view, the radiance level data of O_3 , H_2O and NO and other minor species provided the first measured data which revealed the distribution of these species at high altitude.

The radiance level data of O_3 observed at a tangent height of 80 km show a rather dramatic change from the day to the night side, while the density of the excitation photons remains unchanged between these two periods. A mechanism which causes the diurnal change must be sought in some photo-chemical effect triggered by the ultraviolet photons.

Modification Applied to the BGND Program

In the course of our study, we made an extensive use of four computation programs: the BGND written by Degges of Visidyne, Inc.; the LOWTRAN by Selby ^{11,12} of then AFGL; the FLASH by Blättner of Radiation Research Associates, Inc.; ^{5,6,7} and the FASCODE by Clough *et al* of AFGL. ^{15,16} We applied some modifications to these programs for improvement of the execution efficiency. The operating system of our central site computing facility is designed to accommodate a large number of the remote terminals. Because of the operating system, we found that the program executes more efficiently if its working space required is reduced in size from the original. We divided these programs into small sections, achieving their executions through our remote terminal possible during the regular hour. In addition, inefficient IO operations found in these programs are streamlined to meet our specific needs and to reduce a

computational waste. We have consequently achieved a substantial improvement in running these programs. Our strategy for using the FLASH program was described in our Report No. 1. Overall modifications, except for the BGND, did not extend to rewrite their computation algorithms. We feel that they have no specific merit for their coverage in this report. The case for the BGND is different. Our modification is listed in the Appendix. A general logic flow of the program execution is illustrated in Figure 8.

Other Efforts

- (1) For the balloon-borne experiment, we evaluated the test data on the interferogram PCM telemetry scheme. We found some mysterious bit droppings in the PCM interferogram data string. A substantial effort was poured to resolve this mystery in a joint effort with the Idealab and the AFGL. No conclusions were made on this problem. Our effort was switched after a new contract for the balloon-borne data analysis became effective.
- (2) A very efficient CRT display program for the CYBER system was written by R. Lauzzana. Its listing is given in Appendix B.
- (3) In anticipating a large amount of the CYBER generated plots, we implemented a micro-computer controlled linkage between our plotting machine and the CYBER MODEM scheme. The program implemented on ROM of the 8080 system is listed in Appendix C.

Conclusion

The SPIRE experiment is in a sense a historical event in the progress of atmospheric physics. Even though the data were collected using a simple spectrometer radiometer, their merit is extremely valuable in developing our understanding of the atmospheric radiative process. Our simple-minded model

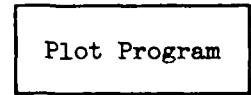
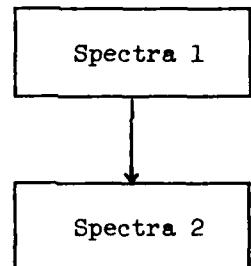
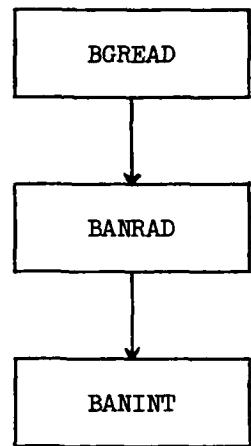


Figure 8

achieved a good success in understanding the radiance level of the major atmospheric CO₂ bands. We believe that our constructed model provides a good insight into the atmospheric radiative process. Ours is no different from the aforementioned computation programs, since all of these are constructed on the same theoretical development. Ours attempted to model the radiative process in the simplest form possible required for providing a good estimate of its level.

There will be no question that the data will be collected with an improved spectral resolution, as time will progress. Then we will come to understand more details of the radiative process in the atmosphere. In so doing, we will be more appreciative of those aforementioned computation programs. At this moment, our simple-minded model functioned reasonably well for analysis of the SPIRE data, revealing a basic character of the infrared radiance level in the upper atmosphere, as well as some abnormal feature which must be studied more carefully in the future.

The following list summarizes our model of the radiative process in the upper atmosphere:

- (1) In the upper atmosphere, the molecular absorption transitions which trigger the emission are well approximated by assuming that they originate from the ground state.
- (2) The photons available for exciting the molecular transitions are either those of the telluric origin or of the solar origin. The cross-over takes place in the vicinity of 2000 cm⁻¹.
- (3) Below the critical altitude, the molecular bands establish a local thermal equilibrium with the photon field. The radiative level given by the blackbody radiation formula is maintained over a large range of altitude if the absorption is very strong. It is independent

of the observation condition and of the scattering radiance level.

- (4) Well above the critical altitude, the radiative level of the molecular band is proportional to the molecular concentration at that altitude in the first-order approximation.
- (5) The scattering radiance level in the near infrared region is primarily determined by the Rayleigh process with a single scattering center.

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Appendix A

```

C C FFCGRAM EGREGAC(OUTPUT,TAPE1,TAPE2,TAPE6)
THIS IS THE FIRST PART OF EGG'S PGNC PROGRAM. IT JUST READS EGG'S
DATA AND GET'S DATA READY FOR PROGRAM BANRAD.
LAST CHANGE AUG 11, 1972
DIMENSION XIC(4,2E),XICT(4),ALT(30),STORE(30),IALT(25)
DIMENSION CATATY(2E),CCEXT(2E),CCEXB(25),NLEVEL(25),NPAND(25)
DIMENSION LSC(25,2E),LCC(25,2E),RCM(25,25),BV(25,25),AMASS(25)
DIMENSION ZANE(7),STAU(7),SFAD(8),RCN(8),RATIC(8)
DIMENSION LEC(25,25),LEL(25,25),STR(25,25),SFLLX(25,25),TFLUX(25,
1E,3)
DIMENSION CC2(18,2),WATVA(12,2)
DIMENSION CHEML(1E1)
REAL LAMCA
  COMM CN /VBRCCM/ I,CC2,WATVA,CCEXT,CCEXB
  COMM CN /TEQCCM/ NLTE
  COMM CN /KCEM/KCHEM
  COMM CN /IFCUT/NR,NF,NW
  COMM CN /CN(201,8),T(201),LAMCA(25),
1 TCTDEN(21)
  DATA(ZAN(I),I=1,7)/E.0.,20.0.,40.0.,60.0.,70.0.,80.0.,90.0/
  DATA(CATATY(I),I=1,25)/EFMFTHANE,8HCAFBN D,EFWATER VA,EHNITRIC C
1, EHNITROLS,EHOZCNE PR,EHCZCNE NC,18*8/
  KCHEM = 1
  NR = 1
  NP = 2
  NTX = 6
  NW = 6
10 FCRMAT (I2)
  READ(NR,10) IGAS
20 FCRMAT (I1,I2) !NUMBER OF GAS FFCFILES, I3)
  WRITE(NW,20) IGAS
  IGAS=IGAS+1
  CC 17[ J=1,IGAS
30 FCRMAT (4AE,I3,F7.0)
  READ(NR,20)(XID(K,J), K=1,4), ICHK, ALTM
  IF(J.EQ.ICHK) GO TC 50
40 FCFMAT (7HICARE, A6, 20HCUT OF CRDER FOR GAS, I3)
  WRITE(NW,40)(XID(K,J),K=1,4), J
50 CC 105 K=1,E
  K1=(K-1)*5+1
  K2=K1+4
60 FCRMAT (E(F5.0,F8.1,2X),I2,I3)
  READ(NR,60)(ALT(L),STCRE(L),L=K1,K2), ICHK, KCHK
  IF(ICFK.EG.J) GO TC 80
70 FCPMAT (3HICDATA CARE CUT OF CRDER FOR GAS,I3, 3X, 8HCARE NO., I3)
  IFIVE = 5
  WRITE(NW,70) J,IFIVE
80 IF(KCHK.EG.K) GO TC 90
  WRITE(NW,70) J,K
90 CC 100 L=K1,K2
  IF(ALT(L).LT.ALTM) GO TC 99
  KSET=L
  GO TO 12[

99 CCNTINUE
100 CCNTINUE
101 CCNTINUE
110 FCRMAT(4E!MAX ALT. NCT FCLNC ON DATA CARDS - GAS NO., I3)

  WRITE(NW,110) J
120 FC 130 L=1,KSET
130 STORE(L)=ALCG(STCRE(L))
  IF(ALTM.GT.150.) GC TC 140
  JALT(J)=ALT-59.
  GC TC 15[

140 IALT(J)=E1.4*(ALTM-150.)/5.
150 CALL INTERP(ALT,STCRE,KSET,1,J)
  L1=ALT(J)
  CC 160 L=1,L1
160 CCNC(L,J)=EXF(CCNC(L,J))
170 CCNTINUE
  CC 175 J=1,201
175 TCTDEN(J)=CCNC(J,IGAS)
  READ(NR,20)(XID(K),K = 1,4), ICHK, KCHK
  IF(ICFK.EG.1) GO TC 19[

180 FCRMAT (7HICARE, 4AE,2EHCLT OF CRDER FOR TEMP DATA)
  WRITE(NW,180)(XICT(K),K=1,4)
190 CC 245 K=1,E
  K1=(K-1)*5+1
  K2=K1+4
  READ(NR,60)(ALT(L),STCRE(L),L=K1,K2), ICHK, KCHK
  IF(ICFK.EG.1) GO TC 21[

200 FCRMAT(31HTEMP DATA CARE CUT OF CRDER NO..T+) 
```

```

10 WRITE(NW,200) K
220 IF(IGAS.EQ.0) GC TC 230
230 CC 240 L=K1,K2
IF(ALT(L).LT.ALTH) GC TC 240
KSET=L
GC TO 260
240 CCNTINUE
245 CCNTINUE
250 FORMAT(37HMAX ALT. NOT FCUNC CN TEMP DATA CARDS)
WRITE(NW,250)
260 CALL INT ERF(ALT,STCRE,KSET,2,1)
261 FCRMAT (1X,IE,1F6E12.4)
WRITE(NW,261) (J,(CCNC(J,I),I=1,IGAS), T(J),J = 1,201)
IGAS=IGAS-1
CC 331 I I=1,IGAS
READ(NR,310) TYPE,NLEV ,NEAN ,AMAS ,CEXT,DEXB
310 FCFMAT (E,22X,I2,3X,I2,3X,FE.2,4X,2(E12.4,3X))
CC 311 J =1,25
IF(TYPE.EQ.CATATY(J)) GC TC 312
311 CCNTINUE
WRITE(NW,370)
370 FCRMAT(3E12.4 THE GAS IS NOT LISTED IN THE PROGRAM)
C STOP 1
312 NSF=J
NLEVEL(NSF)=NLEV
NEAN(NSF)=NEAN
AMASS(NSF)=AMAS
CEEXT(NSF)=CEXT
CEDEXB(NSF)=CEDEXB
IF(NSF.EQ.2)READ(NR,320) (CC2(I,1),CC2(I,2),I=1,18)
320 FCRMAT (E(F5.2,F8.2))
IF(NSF.EQ.3)READ(NR,320) (WATVA(I,1),WATVA(I,2),I=1,18)

ILEVL=0
DC 313 I =1,NLEV
ILEVL=ILEVL+1
READ(NR,330)LSC(IJ,NSF),LCC(IJ,NSF),RCM(IJ,NSF),BV(IJ,NSF),JGAS,
330 FCRMAT (EIE,2E12.4,2EX,2I3)
IF(NSF.EQ.JGAS) GC TC 314
WRITE(NW,340) NSP
340 FCRMAT(64) THE CODE ON THE DATA DOES NOT AGREE WITH THE FIRST CARD
1WFICH IS,I3)
C STOP 2
314 IF(ILEVL.EQ.JLEVEL) GC TC 313
WRITE(NW,350) NSF
350 FCRMAT(42) THERE IS A CARD OUT OF SEQUENCE IN GAS NO.,I3)
C STOP 3
313 CCNTINUE
ILEVL=0
DC 314 IJ=1,NPAN
ILEVL=ILEVL+1
READ(NR,360)LBC(IJ,NSP),LEU(IJ,NSP),STR(IJ,NSP),SFLUX(IJ,NSP),
1(SFLUX(IJ,NSP),J=1,3),JGAS,JLEVEL
STR(IJ,NSP)=2.5793E+10*STR(IJ,NSP)
360 FCPMAT (2IE,1PE12.4,2I3)
362 FORMAT (2TE,1PE12.4,2I3,1X,4FE)
IF(NSF.EQ.JGAS) GC TC 322
WRITE(NW,340) NSF
322 FCRMAT (1F6E12.4)
IF(ILEVL.EQ.JLEVEL) GC TC 316
WRITE(NW,350) NSF
C STOP 4
316 CCNTINUE
317 CCNTINUE
378 FCRMAT (1F6E12.4)
READ(NR,378) CHEML
380 READ(NR,*) IGAS,NLTE
IF(IGAS.EQ.0) STOP
WRITE(NP,378) T
WRITE(NP,378) TOTLEN
WRITE(NP,378) (CCNC(NZ,IGAS),NZ=1,201)
WRITE(NP,360) IGAS,NLTE
NLEV = NLEVEL(IGAS)
NBNC = NEAN(IGAS)
WRITE(NP,335) (XID(K,IGAS),K=1,4),NLEV,NBNC,AMASS(IGAS)
2,CEEXT(IGAS),CEDEXB(IGAS)
335 FCFMAT(4AE,EX,T2,3X,I2,3X,FE.2,4X,2(E12.4,3X))
DC 301 I =1,NLEV
WRITE(NP,360) LSC(IJ,IGAS),LCC(IJ,IGAS),RCM(IJ,IGAS),BV(IJ,IGAS)
381 CCNTINUE

```

```

CC 382 IJ = 1, NEND
      WRITE(NE,360) LEC(IJ,IGAS),BU(IJ,IGAS),STR(IJ,IGAS)
 2 ,SFLLX(IJ,IGAS),TFLUX(IJ,IGAS,1)
382 CONTINUE
      JALT = (JALT(IGAS)-1)/5
      WRITE(NF,360) JALT(IGAS),JALT
      IF(IGAS.NE.4) GO TO 830
      WRITE(6,378) CHEML
      GC TO 830

ENC
      SLBROLTIME INTERP(ALT,STORE,KSET,MODE,I)
      C1MENSICK ALT(30),STORE(30)
      CCMMCN CCNC(1201,8),T(201),LAMEA(25),
      1TCDE(2(1))
      J=1
      A=60
 8   CC 120 K=1,KSET
10   IF(A-ALT(K))80,20,20
20   GC TO (30,40),MODE
30   CCNC(J,I)=STORE(K)
      GC TO 50
40   T(J)=STORE(K)
50   IF(K.LT.KSET) GO TO 60
      RETURN
60   IF(J.GE.51) GO TO 70
      A=A+1.
      GC TO 110
70   A=A+5.
      GC TO 110
80   FR=(A-ALT(K-1))/(ALT(K)-ALT(K-1))
      GC TO (90,100),MCCE
90   CCNC(J,I)=STORE(K-1)+FR*(STORE(K)-STORE(K-1))
91   J=J+1
      IF(J.GT.51) GO TO 92
      A=A+1.
      GC TO 10
92   A=A+5.
      GC TO 10
100  T(J)=STORE(K-1)+FR*(STORE(K)-STORE(K-1))
      GC TO 91
110  J=J+1
120  CCNTINUE
      RETURN
ENC

```

```

      FRCGRAMP FAFAC(TAFFE2=103E,TAFF4=103B,T/PE3,OUTPUT=103E
      2 ,TAFF6=CLTFLT)
      THIS IS A MODIFIED VERSION OF CECG'S SUBROUTINE BANRAC
      LAST CHANGE AUG 30, 1978
      DIMENSION FCWER(40),KL(40),XIC(40)
      DIMENSION TFLUX(40),SFLLX(40)
      DIMENSION CFAC(40)
      DIMENSION LSC(40),LCC(40),LPU(40),LPC(40)
      DIMENSION WAVE(40),STR(40),EV(40),ANCLAM(40),FACC(40)
      DIMENSION GNESTE(201),TEMF(201),CENTCT(201),VIE(40,3)
      DIMENSION ENCRAD(201,40),ENETCT(201,40)
      CCMCRA /VIEFCP(201,25)
      CCMCRA /CCL /TVSIG(201),VTSIG(201),VVLSIG(201),VVCSIC(201)
      2 ,VVNSIG(201),VVRSIG(201)
      CCMCRA /CFCAFA /NALT,NLEVEL,NEAND,ZFAC,KP(40),EA(40),EU(40)
      2 ,GCCWN(40),TTUP(40),TTECHN(40),LFSTA(40),DOWN(40),TTF(201,40)
      3 ,CCNC(201),CFEML(201)
      EQUIVALENCE (GNESTE(1),VIRFCF(1,1))
      EGLIVALENCE (TFLUX,GLF),(SFLLX,GECWN)
      EGLIVALENCE (TTF,ENCTAL),(TTF,ENCRAD)
      EQUIVALENCE (LSC(1),VIE(1,1)),(LCC(1),VIE(1,2)),(LBU(1),VIE(1,3))
      EQUIVALENCE (CFAC,CCWN),(CENTCT(1),TTF(1,1))
      EXTERNAL CC2,H2C,FAC,C3
      ZFAC=1.0
      C READ IN TEMPERATURE AND MOLECULAR PROFILE
      READ(2,378) TEMP
      WRITE(14,378) TEMP
      READ(2,378) DENTCT
      READ(2,378) CCNC
      378 FCFMAT(1FE12.4)
      C READ IN WHICH GAS AND CONCENTRATION (NLTF=3 NIGHT, NLTE=4 DAY)
      READ(2,368) IGAS,NLTE
      WRITE(14,368) IGAS,NLTE
      368 FCFMAT(3E12.4)
      C NOTE I IN LEGS IS ICAS IN THIS PROGRAM
      C READ IN LEVEL INFO
      READ(2,310) (XIC(K),K=1,4),NLEVEL,NEAND,RM,CDEXT,CCEXE
      WRITE(14,310) (XIC(K),K=1,4),NLEVEL,NEAND,RM,CDEXT,CCEXE
      310 FCFMAT(4E12.4),EX,I2,2X,I2,3X,FE.2,4X.2(1E12.4,3X))
      CC 1 I=1,NLEVEL
      READ(2,368) LSC(I),LCC(I),WAVE(I),PV(I)
      WRITE(14,368) LSC(I),LCC(I),WAVE(I),PV(I)
      C READ EAND INFO
      CC 2 I=1,NEAND
      2 READ(2,368) LEC(I),LBU(I),STR(I),SFLLX(I),TFLUX(I)
      READ(2,368) NALT,ALT
      IF(IGAS.EQ.4) READ(2,50) (CFEML(I),I=1,1E1)
      50 FCFMAT(5E12.4)
      C THIS SECTION COMPUTES POPULATION OF VIBRATIONAL LEVELS.
      C COMPUTE THERMAL POPULATION
      CC 130 N7 = 1, NALT
      SUM = 0.0
      TMP = TEMF(NZ)
      CC 110 NL = 2, NLEVEL
      XYN = MCC(LSC(NL)/10,10)
      AA = EXP(-1.43875*WAVE(NL)/TMP)*XYN
      VIBFCP(NZ,NL) = AA
      110 SUM = SLM + AA
      CCNTINUE
      SUM = SLM + 1.0
      GNESTE(NZ) = CCNC(NZ)/SLM
      VIBFCP(NZ,NL) = CCNC(NZ)*VIBFCP(NZ,NL)/SUM
      120 CCNTINUE
      130 CCNTINUE
      C FIND LEVEL COUPLED WITH NITROGEN
      CC 135 NL = 2, NLEVEL
      IF (LCC(NL).EQ. 1) VVWAVE = 2731. - WAVE(NL)
      IF (LCC(NL).EQ. 2) VVWAVE = 1556.4 - WAVE(NL)
      FAC = 2331.0
      IF (LCC(NL).EQ. 2) FAC = 1556.4
      135 CCNTINUE
      IF(NLTE.EQ.1) GO TO 170
      C FIND COLLISION EXCITATION AND DE-EXCITATION COEF'S
      CC 160 NZ = 1,NALT
      TEMPNZ = TEMP(NZ)
      TEMF13 = EXP(ALOG(TEMPNZ)/3.0)
      IF (IGAS.EQ. 3) GO TO 136
      IF (IGAS.NE. 2) GO TO 140
      VTSIG(NZ) = 6.69E-10*EXP(-84.07/TEMP13)
      VVRSIG(NZ) = 1.71E-6*EXP(-172.2/TEMP13)
      1 + E.07E-14*EXP(15.27/TEMP13)

```

$VVDSIG(NZ) = 1.0E-15 + 5.1EE-11 * EXP(-70.75/TEMP15)$
 $VVNSIG(NZ) = VVRSIG(NZ) / EXP(1.43879/TEMP(NZ) + 18.0)$
 $VVUSIG(NZ) = VVDSIG(NZ) * VIEFCF(NZ, 8) / VIEPOP(NZ, 5)$
 GO TC 150
 130 CONTINUE
 $VVNSIG(NZ) = 4.0E-13$
 $VTSIG(NZ) = 5.37E-16 * EXP(-70.0/TEMP13)$
 $VVRSIG(NZ) = VVNSIG(NZ) * EXP(-1.43879/TEMP(NZ) + 1E94.736)$
 GO TC 150
 140 CONTINUE
 $VTSIG(NZ) = CDEXT$
 $VVNSIG(NZ) = CDEXE$
 $VVRSIG(NZ) = VVNSIG(NZ) * EXP(-1.43879/TEMP(NZ) * VVWAVE)$
 150 CONTINUE
 $VVSIG(NZ) = VTSIG(NZ) / GNCSTE(NZ) * VIEPCP(NZ, 2)$
 $VTSIG(NZ) = VTSIG(NZ) * CENTOT(NZ)$
 $VVSIG(NZ) = VTSIG(NZ) * CENTCT(NZ)$
 $TNNV = TEMF(NZ)$
 IF (ALTE .LT. 5) GO TC 154
 IF (NZ .GT. 41 .AND. NZ .LT. 51) TNNV = TEMP(NZ) + FLCAT(NZ-41)
 1 * 280.
 IF (NZ .GE. 51) TNNV = 3000.0
 154 CONTINUE
 $FACTCR = EXP(-1.43879 * FAC / TNNV)$
 $VVNSIG(NZ) = VVNSIG(NZ) * CENTCT(NZ) * FACTOR / (1.0 + FACTCR)$
 $VVRSIG(NZ) = VVRSIG(NZ) * CENTCT(NZ) / (1.0 + FACTOR)$
 160 CONTINUE
 C COMPUTE VIEFCP FCF TEMF=180
 DC 160 NZ = 1, NALT
 $SUM = 0.0$
 $TMP = 1E0.0$
 DC 162 NL = 2, NLEVEL
 $XVN = MO(LSC(NL)) / 10.10$
 $AA = EXP(-1.43879 * WAVE(NL) / TMP) * XVN$
 $VIBFCP(NZ, NL) = AA$
 $SUM = SLM + AA$
 162 CONTINUE
 $SUM = SLM + 1.0$
 $GNDSTE(NZ) = CCNC(NZ) / SUM$
 DC 164 NL = 2, NLEVEL
 $VIRFCF(NZ, NL) = (CCNC(NZ) * VIEFCP(NZ, NL)) / SUM$
 164 CONTINUE
 166 CONTINUE
 170 CONTINUE
 C COMPUTE CFTICAL DATA
 DC 170 NE = 1, NEFAC
 $KBC = LEC(NB)$
 $NI = MOC(NEC/100, 100)$
 $NF = MOC(NEC, 100)$
 $KL(NE) = MCC(LBU(NB)) / 10, 10$
 $KB(NE) = MCC(LBU(NE)), 1C$
 $TWAVE = WAVE(NI) - WAVE(NF)$
 $ENCLAM(NB) = 1.0E+4 / TWAVE$
 1 IF (TFLLX(NB) = EG, 0, 1) TFLLX(NB) =
 1 $5.957E+6 * TWAVE^{+4} / (EXP(1.43879 * TWAVE / TEMP(1)) - 1.0)$
 $XYM = MCC(LSC(NF)) / 10, 10$
 $XYMM = MCC(LSC(NI)) / 10, 10$
 $EA(NE) = STR(NB) * 2.804 / ENCLAM(NE) * 2 * XYMM$
 1 $* EXP(1.43879 * WAVE(NF) / 296.0) / XYMM$
 $FCWER(NE) = 1.580E-5E-20 / ENCLAM(NB)$
 $GFAC(NB) = 3.72E-24 * STR(NB) * ENCLAM(NB) * 2$
 1 $* EXP(1.43879 * WAVE(NF) / 296.0)$
 $CUP(NB) = GFAC(NE) * TFLLX(NB) * 12.4864$
 IF (ALTE .LT. 3) GLF(NB) = 0.0
 $GCOWN(NE) = GFAC(NE) * SFLUX(NE)$
 IF (ALTE .LT. 4) GCOWN(NB) = 0.0
 $KBN = KE(NE)$
 $KLN = KL(NE)$
 $EVN = BV(1)$
 $ENDL = ENCLAM(NB)$
 $FACC(NB) = FCWER(NB) * EA(NB)$
 $SNB = STR(NB) * EXP(1.43879 * WAVE(NF) / 296.0)$
 DC 170 NZ = 1, NALT
 $ENDTAU(NZ, NB) = TAU(NB, KEN, RM, BVN, TEMF(NZ), BNCL, SNB,$
 1 VIBFOP(NZ, NF))
 IF (ALTE .GT. 1) GO TC 180
 $ENDFAC(NZ, NB) = VIEFCF(NZ, NI) * FAC$
 180 CONTINUE
 DC 175 NE = 1, NEFAC
 WRITE(4, 360) LBC(NE), LEL(NB), STR(NB), ENCLAM(NB), FACC(NB)
 WRITE(4, 360) NALT, JALT

```

        IF (NLTE .EQ. 1) GO TC 900
        ITMAX = 5
        IF (IGAS .EQ. 2) ITMAX = 10
        AC = .3
        ITER = 0
C      START OF BALANCE LCCF
181     ITER = ITER + 1
        NMIN = NALT - 1
C      CC 190 NB = 1, NBAND
C      LCCF CUT ENCTAU AND TTH ARE EQUIVALENT
        TTHL = SIMF(ENCTAU(1,NE),1.0E+5,1)
        CC 182 NZ=2,NMIN
        NZM=NZ-1
        TTHNZ = SIMF(ENCTAU(NZM,NE),1.0E+5,2)
        TTH(NZM,NB) = TTHL
        TTHL = TTHNZ
182     CCNTINUE
        TTH(NMIN,NE) = TTHL
        SUM = 0.0
        IF (NLTE .LE. 3) GO TC 186
        CC 184 NZ=1,NMIN
184     SLM = SLM + TTH(NZ,NE)
        TTDOWN(NE) = SUM
        TTUP(NB) = 0.0
        DCWN(NB) = SUM
        UFSTA(NB) = 0.0
190     CCNTINUE
200     CCNTINUE
C      NIAL SHOULD BE A FOWER OF TWO
        NIAL = 4
        WRITE(3,1000)
1000    FORMAT(* VIBFOP *)
        CC 1020 NZ = 1, 201
1020    WRITE(3,8) (VIBPOF(NZ,NL),NL=1,NLEVEL)
        WRITE(6,8) (VIBPOF(12,NL),NL=1,NLEVEL)
        WRITE(6,8) (VIBPOF(32,NL),NL=1,NLEVEL)
        WRITE(3,1001)
1001    FCRRMAT(1, NALT,NBAND,ZFAC,KB(4C),EA(4))
        WRITE(3,*) NALT,NEANC,7FAC
        WRITE(3,*) (KE(NB),NE=1,NEANC)
        WRITE(3,1008) (EA(ICUT),ICUT=1,16)
1008    FCRRMAT(1,FPE12,4)
        GO TC (210,220,230,240,250,260,270,280), IGAS
210     CONTINUE
        GO TC 350
220     CCNTINUE
        CALL CAL21(CC2,NIAL,AC,VIE)
        GO TC 350
230     CCNTINUE
        CALL CAL2(F2C,NIAL,AC,VIE)
        GO TC 350
240     CCNTINUE
        CALL CAL2(FNO,NIAL,AC,VIE)
        GO TC 350
250     CCNTINUE
        GO TC 350
260     CCNTINUE
        CALL CAL2(C3,NIAL,AC,VIE)
        GO TC 350
270     CCNTINUE
        CC 360 NE = 1, NBAND
        NBC = LEC(NB)
        NF = MOL(NBC,100)
        KLN = KL(NE)
        KEN = KE(NE)
        RVN = BV(1)
        SNB = STR(NB)*EXF(1.43879*WAVE(NF)/29E.0)
        BRDL = BICLAP(NB)
        CC 360 NZ = 1, NALT
        ENDTAU(NZ,NE) = TAUMAX(KLN,KEN,RM,RVN,TEMP(NZ),BNCL,SNB,
1        VIBFOP(NZ,NF))
360     CCNTINUE
400     CCNTINUE
        IF (ITFF .LT. ITMAX) GO TC 181
500     CCNTINUE
        IF (NLTE .LE. 2) GO TC 920
        CC 91(NZ = 1, NALT
        WRITE(4,8) (VIBFCP(NZ,NL), NL = 1, NLEVEL)
510     CCNTINUE
        CC 91(NZ = 1, NALT
        WRITE(4,8) (ENCTAU(NZ,NE),NE=1,NBAND)
520     CCNTINUE

```

8 STCP
FORMAT (1F8E12.4)
END

C
C SLBFOLTIPE CALZ(GAS,NIAL,AC,VIE)
THIS SUBROUTINE CALCULATES VIE FCP FOR A GIVEN GAS AT ALL ALTITUDES
CIMNSICK VIE(40,3)
FCMMCN /CFCATA/ NALT, NLEVEL, NBEAD, ZFAC, KB(40), EA(40), GUF(40)
2 ,GDOWN(40), TTUP(40), TTECHN(40), UPSTA(40), DOWN(40), TTF(201,40)
3 ,CONC(201), CHEML(201)
EXTERNAL GAS
WRITE(6,*) NZ,NIAL,AC
NZ=1
10 CALL CALVIE(GAS,NZ,NIAL,NEXT,AC,VIR)
NZ=2*NIAL+NZ
NIAL=NEXT
IF((NZ+2*NEXT).GT.NALT) NIAL=(NALT-NZ)/2
IF(NIAL.LE.0) RETLN
GO TO 10
END

C
C SLBFOLTTNE CALVIR(GAS,NZST,NIAL,NEXT,AC,VIP)
THIS SUBROUTINE CALCULATES VIE FCP FOR THREE ALTITUDES AND INTERPOLA
BETWEEN THE END POINTS TO CHECK ACCURACY AND CHANGFS THE ALTITUDE
INTERVAL IF NECESSARY
NCTE: UPSTA,DCWN,TTUP,TTECHN MUST BE SET BEFORE CALLING THIS RCLTNE
CIMENSION VIE(40,3),ATNL(10)
CCMMCN VIEFCF(201,25)
CCMMCN /CFCATA/ NALT, NLEVEL, NBEAD, ZFAC, KB(40), EA(40), GUF(40)
2 ,GDOWN(40), TTUP(40), TTECHN(40), UPSTA(40), DOWN(40), TTF(201,40)
3 ,CONC(201), CHEML(201)
EXTERNAL GAS
WRITE(3, 100) NZST,NIAL,NEXT,AC
FCRMAT(*,CALVIB HERE WITH NZST=*,I5,* NIAL=*,IE,* NEXT=*,I5
2 ,* AC=*,F5.3)
IF(NZST.EQ.1) CALL GAS(VIE(1,3),NZST)
CC 10 NL=1,NLEVEL
10 VIB(NL,1)=VIE(NL,3)
CALL TTSET(NZST,NIAL,1)
NZ=NZST+NIAL
CALL GAS(VIE(1,2),NZ)
CALL TTSET(NZ,NIAL,2)
NZ=NZ+NIAL
CALL GAS(VIE(1,3),NZ)
CHECK INTERF VALUES
SUM=0.
CC 30 NL=1,NLEVEL
ATNL(NL)=ABS(ABS(VIE(NL,1)+VIE(NL,3))* .5-ABS(VIE(NL,2)))
30 SLM=ATNL(NL)+SUM
ATEST=SUM/(FLCAT(NLEVEL))
WRITE(3, 150) ATEST
150 FCRMAT(*,ATEST=*,F10.3)
IF(ATEST.LE.AC) GO TO 50
C REDUCE ALTITUDE RANGE IF POSSIBLE
NT=NIAL/2
RT=FLCAT(NIAL)/2.
IF(FLCAT(NT).NE.RT) GO TO 50
NIAL=NT
CC 40 NL=1,NLEVEL
40 VIB(NL,3)=VIE(NL,2)
CALL TTSET(NZST,NIAL,1)
NZ=NZST+NIAL
CALL GAS(VIE(1,2),NZ)
GO TO 20
C QUADRATIC FIT
50 STEP=NIAL*.2
STEP2=STEP*STEP/2.
CC 60 NL=1,NLEVEL
AFIT=(VIE(NL,1)+VIE(NL,3)-2*VIE(NL,2))/STEP2
BFIT=(VIE(NL,3)-VIE(NL,1))/STEP2
CFIT=VIE(NL,2)
XMD=NZST+NIAL
NZL=2*NIAL+NZST
CC 60 NZ=NZST,NZL
X=NZ-XMD
C THE VIB FCF IN VIE(I,J) IS LCG OF DENSITY
60 VIBPOF(NZ,NL)=EXP((AFIT*X+BFIT)*X+CFIT)
C RESET UPSTA, DOWN
CALL TTSET(NZST,2*NIAL,1)

```

CC 70 NB=1,NEAND
UFSTA(NB)=TTUF(NB)
C CWN(NB)=TTCWN(NB)
NEXT=NIAL
C HAS SAFETY FACTOR AGAINST OSCILLATION OF 2.
IF((ATES1*4.)<LT.AC) NEXT=NEXT*2
RETURN
END

C SLRRCLTINE TTSET(NZST,NIAL,PCC)
C MCC=1 IS TAU FROM UFSTA AND CCWN
MCF=2 TAU FROM TTUP AND TTCWN
CCMMOK /CFCATA/ MALT,NLEVEL,NEAND,ZFAC,KB(40),EA(40),GUF(40)
2 ,GDCWN(40),TTUP(40),TTCWN(40),UFSTA(40),DOWN(40),TTH(201,40)
3 ,CONC(201),CHEML(201)
IF(MCC.NE.1) GO TO 20
CC 10 NB=1,NEAND
TTUP(NB)=UFSTA(NB)
10 TTDOWN(NB)=CCWN(NB)
NZF=NZST+NIAL-1
CC 30 NZ=NZST,NZF
CC 30 NB=1,NEAND
TA=TTH(NZ,NE)
TTUP(NB)=TTUP(NB)+TA
20 TTDOWN(NB)=TTCWN(NE)-TA
RETURN
END

C FUNCTION RTFUN(NZ,NUS,NES)
C NUS IS THE NUMBER OF THE LEVEL, NPS IS THE NUMBER OF THE BAND
DIMENSION GNDSTE(201)
CCMMOK VIEFCF(201,25)
CCMMOK /CFCATA/ MALT,NLEVEL,NEAND,ZFAC,KB(40),EA(40),GUF(40)
2 ,GDCWN(40),TTUP(40),TTCWN(40),UFSTA(40),DOWN(40),TTH(201,40)
3 ,CONC(201),CHEML(201)
EQUIVALENCE (GNDSTE(1),VIBFCF(1,1))
RTFUN=
1 (XFER(VIBFOP(1,NUS),TTH(1,NFSI),NZ,1,1,KB(NES))
2 + XFER(VTEFOP(NZ,NUS),TTH(NZ,NES),MALT,NZ,2,KB(NES)))*EA(NBS)
3 + GDCWN(NFSI)*SM1(KB(NES),ZFAC*TTCWN(NPS))
4 + GUF(NES)*SN12(KB(NES),TTUF(NPS))
RETURN
END

SLRRCLTINE CC2(VIEI,NZ)
DIMENSION VIEI(40),GNDSTE(201)
CCMMOK /CCL/ TVSIG(201),VTSIG(201),VVUSIG(201),VVDSIG(201)
2 ,VVNSIG(201),VVRSIG(201)
CCMMOK /VIEFCF(201,25)
CCMMOK /CFCATA/ MALT,NLEVEL,NEAND,ZFAC,KB(40),EA(40),GUF(40)
2 ,GDCWN(40),TTUP(40),TTCWN(40),UFSTA(40),DOWN(40),TTH(201,40)
3 ,CONC(201),CHEML(201)
EQUIVALENCE (GNDSTE(1),VIEFCF(1,1))
VIRI(8)=ALCG(VNDSTE(8))+VVNSIG(NZ)+RTFUN(NZ,P,S)
2 /(VVRSIG(NZ)+EA(9)+VVDSIG(NZ))
TVS=VTSIG(NZ)
VTS=VTSIG(NZ)
FOP3=VIEFCP(NZ,3)
FOP4=VIEFCP(NZ,4)
FCP7=VIEFOP(NZ,7)
FOP2=((TVS+RTFUN(NZ,2,1))*GNDSTE(NZ)
2 + FCP3*(EA(2)+VTS)+FOP4*(EA(3)+VTS)
3 + FCP7*(EA(7)+VTS))/(VTS+EA(1)+TVS)
VIRI(2)=ALCG(POF2)
FOP8=VIEFCF(NZ,8)
FCP7=((TVS+0.25*GLF(7)*SN12(KB(7)),TTUP(7))
1 + GDOWN(7)*SM1(KB(7),TTCWN(7))*FCF2
2 + EA(8)*FCP8+EA(15)*VIEFCF(NZ,9)+EA(16)*VIBFCP(NZ,10)
3 /(EA(7)+VTS)
VIEI(7)=ALCG(POF7)
FCP5=VIEFCP(NZ,5)
FCP6=VIEFCP(NZ,6)
FOP3=(FCF2*(0.25*TVS+GUP(2)*SN12(KB(2),TTLP(2)))
1 + GDOWN(2)*SM1(KB(2),TTCWN(2))+FCF5*(EA(4)+VTS)
2 + EA(10)*FCP4+EA(13)*VIEFCF(NZ,9)+EA(14)*VIBFCP(NZ,10)
3 /(EA(2)+VTS+TVS)
VIEI(3)=ALCG(FOP3)
FOP4=(FCF2*(0.5*TVS+GUP(3)*SN12(KB(3),TTLF(3)))
1 + GDOWN(3)*SM1(KB(3),TTCWN(3))+FCP5*EA(5)
2 + FCP6*(EA(6)+VTS))/(EA(3)+VTS+TVS)

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VIEI(4)=ALCG(FOP4)
FOP5 = (FCF4*(TVS + GLP(4)*SN12(KE(4),TTUF(4)))
1 + GCONN(4)*SM1(KE(4),TTCWN(4)))
2 + FCP4*(GUP(5)*SN12(KP(5),TTUP(5))
3 + GCONN(5)*SM1(KE(5),TTCWN(5))) + FOP8*VVDSIG(NZ)
4 / (EE(4) + EA(5) + VTS + VVUSIG(NZ))

VIEI(5)=ALCG(FOP5)
FCP6 = (FCF4*(0.5*TVS + GUP(6)*SN12(KP(6),TTUF(6)))
1 + GCONN(6)*SM1(KE(6),TTCWN(6)))./ (EA(6) + VTS)

VIEI(6)=ALCG(FOP6)
VIBI(9)=(VIEFCF(NZ,3)*(VNSIG(NZ) + RTFUN(NZ,9,13))
1 + GNDSTE(NZ)*RTFUN(NZ,9,13) + VIEPCP(NZ,7)*RTFUN(NZ,9,15))
2 / (VVRSIG(NZ) + EA(11) + EA(13) + EA(15))
VIBI(10)=ALCG((VIEFCP(NZ,7)*(VNSIG(NZ) + RTFLN(NZ,10,16))
1 + VIBFCF(NZ,3)*RTFUN(NZ,10,14) + GNDSTE(NZ)*RTFUN(NZ,10,12))
2 / (VVRSIG(NZ) + EA(12) + EA(14) + EA(16)))
VIBI(11)=CONC(NZ) - (FOP8+FOP6+FOP5+FOP7+FOP4
1 + PCP3 + FCF2)
IF (VIBI(1).LE. 0.0) VIDI(1) = 0.05*CONC(NZ)

VIDI(1)=ALCG(VIBI(1))
RETURN
END

```

C

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SUBROUTINE H2C(VIEI,NZ)
DIMENSION VIEI(40),GNDSTE(201)
CCMCA VIEPCP(201,25)
CCMCA /CCL/ TVSIG(201),VTSIG(201),VVUSIG(201),VVDSIG(201)
2 ,VVNSIG(201),VVRSIG(201)
CCMCA /CCFATA/ NALT,NLEVEL,NEAND,ZFAC,KB(40),EA(40),GU(40)
2 ,GDCWN(40),TTUP(40),TTCWN(40),UPSTA(40),COWN(40),TTH(201,40)
3 ,CONC(201),CHEML(201)
EQUIVALENCE (GNDSTE(1),VIBFCF(1,1))
VIBI(2)=ALCG((TVSIG(NZ) + VVNSIG(NZ)
1 + RTFUN(NZ,2,1)*GNDSTE(NZ) + VIEFCF(NZ,3)*EA(15))
2 / (VTSIG(NZ) + EA(1) + VVNSIG(NZ)))
VIBI(3)=ALCG((TVSIG(NZ) + VVNSIG(NZ)
1 + RTFUN(NZ,3,5))*VIEFCF(NZ,2) +
2 RTFLN(NZ,3,21)*GNDSTE(NZ) + VIEPCP(NZ,6)*EA(11))
3 / (EA(2) + EA(5) + VTSIG(NZ) + VVNSIG(NZ))
VIBI(4)=ALCG((RTFUN(NZ,4,3)*GNDSTE(NZ)
1 + RTFUN(NZ,4,6)*GNDSTE(NZ))/(EA(3) + EA(6)))
VIBI(5)=ALCG((RTFUN(NZ,5,4)*GNDSTE(NZ)
1 + RTFLN(NZ,5,7)*VIBFCP(NZ,2))/(EA(4) + EA(7)))
VIBI(6)=ALCG((RTFUN(NZ,6,8)*VIBFCF(NZ,8)
1 + RTFUN(NZ,6,12)*GNDSTE(NZ))/(EA(8) + EA(11))
2 + EA(12)))
VIBI(7)=ALCG((RTFLN(NZ,7,9)*VIEFOP(NZ,2)
1 + RTFUN(NZ,7,13)*GNDSTE(NZ))/(EA(9) + EA(13)))
VIBI(8)=ALCG((RTFUN(NZ,8,10)*VIEFOP(NZ,2)
1 + RTFUN(NZ,8,14)*GNDSTE(NZ))/(EA(10) + EA(14)))
VIDI(1)=ALCG(CONC(NZ) - (VIEFCF(NZ,8)*VIEPCP(NZ,7) + VIBFCF(NZ,6)
2 + VIEFOP(NZ,5) + VIEFCF(NZ,4)*VIEFCF(NZ,3) + VIBFCF(NZ,2)))
RETURN
END

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```

SLBRCLTNE RNC(VIEI,NZ)
DIFFREICK VIEI(40),GNDSTE(201)
CCMCA VIEFCF(201,25)
CCMCA /CCL/ TVSIG(201),VTSIG(201),VVUSIG(201),VVDSIG(201)
2 ,VVRSIG(201),VVRSIG(201)
CCMCA /CCFATA/ NALT,NLEVEL,NEAND,ZFAC,KB(40),EA(40),GU(40)
2 ,GDCWN(40),TTUP(40),TTCWN(40),UPSTA(40),COWN(40),TTH(201,40)
3 ,CONC(201),CHEML(201)
EQUIVALENCE (GNDSTE(1),VIEPCF(1,1))
VIBI(2)=ALCG((TVSIG(NZ) + VVNSIG(NZ)
1 + GLP(1)*SN12(KE(1),TTLF(1))
2 + (XFER(VIEPOP(1,2),TTF(1,1),NZ,1,1,KB(1))
3 + XFER(VIEPCP(NZ,2),TTF(NZ,1),NALT,NZ,2,KB(1)))*EA(1)
4 + GCONN(1)*SM1(KE(1),TTCWN(1))*GNDSTE(NZ)
5 + (EA(3) + VTSIG(NZ))*VIEFCF(NZ,3) + 2.0*CHEML(NZ))
6 / (EA(1) + VTSIG(NZ) + VVRSIG(NZ) + TVSIG(NZ)))

```

0

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VIBI(3)=ALCG((TVSIG(NZ)
1 + GLP(3)*SN12(KE(3),TTLF(3))
2 + GLP(2)*SN12(KE(2),TTLF(2))
3 + GCONN(2)*SN12(KE(2),TTCWN(2)))
4 + GCONN(3)*SN12(KE(3),TTCWN(3)))*VIEPCP(NZ,2)
5 + 2.0*(CHEML(NZ))/(EA(2) + EA(3) + VTSIG(NZ)))

```

0 VIBI(1)=ALCG(CONC(NZ) - VIEFCF(NZ,2) - VIEFOP(NZ,3))

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C
      SUBROUTINE Q3(VIBI,NZ)
      DIMENSION VIBI(40),GRCSIE(201)
      CCMMCN VIEFCF(201,25)
      CCMMCN /CCL/ TVSIG(201),VTSIG(201),VVUSIG(201),VVDSIG(201)
      2 ,VVRSIG(201),VVRSIG(201)
      CCMMCN /CFALTA/ RALT,NLEVEL,NEAND,ZFAC,KB(40),EA(40),GU(40)
      2 ,GDCWN(40),TTUP(40),TTECWN(40),UPSTA(40),DOWN(40),TTF(201,40)
      3 ,CONC(201),CHEML(201)
      EQUIVALENCE (GNCSIE(1),VIEFCF(1,1))
      VIBI(2)=ALCG((TVSIG(NZ)+VVNSIG(NZ))
      1 + GLP(1)*SN12(KE(1),TTUF(1))
      2 + XFEF(VIBPOP(1,2),TTF(1,1),NZ,1,1,KB(1))
      3 + XFER(VIEFCP(NZ,2),TTH(NZ,1),RALT,NZ,2,KB(1))*EA(1)
      4 + GDOWN(1)*SM1(KE(1),TTECWN(1))*GRCSIE(NZ)
      5 /(VTSI((NZ)+EA(1)+VVRSIG(NZ)))
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    IF (T .GT. 1.0)  GO TO 110
    X = (T - 0.5)*2.0
    SL03 = (((2.583E-4*X - 6.0149E-3)*X - 0.1027223)*X
    1 + 0.12528821*ALCG(1)*T + (0.4035503)*T
    RETURN
110  CONTINUE
    IF (T .GT. 4.0)  GO TO 120
    X = (T - 2.5)/1.5
    SL03 = (((((7.886E-4 - 3.516E-4*X)*X - 1.0464E-3)*X
    1 + 2.7618E-3)*X - 7.96629E-3)*X + 0.0219502)*T
    2 - 0.06265501*X + 0.157702)*T
    RETURN
120  CONTINUE
    IF (T .GT. 10.0)  GO TO 130
    X = (T - 7.0)/3.0
    SL03 = ((((((1.7495E-4*X - 4.9991E-4)*X + 0.00113772)*X
    1 - 0.00234638)*X + 0.010099E11)*X - 0.03090774)*X
    2 + 0.1028E302)*T
    RETURN
130  CONTINUE
    Y = ALCG(T)
    IF (T .GT. 100.0)  GO TO 140
    X = 5.21034037197E/Y - 3.0
    SL03 = (((((2.93E-5 - 5.78E-5*X)*X + 2.458E-4)*X
    1 - 0.0012858)*X + 0.0340297)*X + 0.2210431)*Y
    RETURN
140  CONTINUE
    IF (T .GT. 1000.0)  GO TO 150
    X = 27.631021115E3/Y - 5.0
    SL03 = (((((8.3E-6 - 2.01E-5*X)*X - 1.843E-4)*X
    1 + 0.0129309)*X + 0.2725344)*Y
    RETURN
150  CONTINUE
    X = 13.6155105579E/Y - 1.0
    SL03 = (((((5.39E-4*X - 0.0(15038)*X + 0.0016309)*X
    1 - 0.0001371)*X - 0.0(23757)*X + 0.0(301341)*X
    2 + 0.23158E0)*Y
    RETURN
200  CONTINUE
    IF (T .GT. 1.0)  GO TO 210
    X = (T - 0.5)*2.0
    SL03 = (((((1.607E-4*X - 4.05E9E-3)*X - 0.08808851)*X
    1 + 0.097407381*ALCG(1)*T + 0.4161293)*T
    RETURN
210  CONTINUE
    IF (T .GT. 4.0)  GO TO 220
    X = (T - 2.5)/1.5
    SL03 = (((((5.904E-4 - 2.794E-4*X)*X - 8.0E9E-4)*X
    1 + 2.2264E-3)*X - 6.6E04E-3)*X + 0.0197235)*X
    2 - 0.06315E2)*X + 0.2E07132)*T
    RETURN
220  CONTINUE
    IF (T .GT. 10.0)  GO TO 230
    X = (T - 7.0)/3.0
    SL03 = (((((1.478E-4*X - 4.291E-4)*X + 1.0472E-3)*X
    1 - 3.316E-3)*X + 0.0107584)*X - 0.03E6259)*X
    2 + 0.1310340)*T
    RETURN
230  CONTINUE
    Y = ALCG(1)
    IF (T .GT. 100.0)  GO TO 240
    X = 5.21034037197E/Y - 3.0
    SL03 = (((((4.188E-4 - 2.131E-4*X + 3.79E-5)*X)*X
    1 + 7.79E-5)*X + 5.04E-5)*X - 0.0226E9)*X
    2 + 0.03208891)*X + 0.41E7794)*Y
    RETURN
240  CONTINUE
    IF (T .GT. 1000.0)  GO TO 250
    X = 27.631021115E3/Y - 5.0
    SL03 = (((((1.65E-5*X - 2.842E-4)*X + 0.01347E7)*X
    1 + 0.3700E89)*Y
    RETURN
250  CONTINUE
    X = 13.6155105579E/Y - 1.0
    SL03 = (((((1.0049E-3 - 4.412E-4*X)*X - 8.441E-4)*X
    1 + 9.451E-4)*X - 3.4768E-3)*X + 0.0235987)*X
    2 + 0.3254546)*Y
    RETURN
300  CONTINUE
    SL03 = 0.0
    RETURN
END

```

```

FUNCTION SN12 (K,TAL)
T = ABS(TAL)
IF (T .EQ. 0.0)  GC TC 300
IF (K .LT. 1)  GC TC 200
IF (T .GT. 1.0)  GC TC 110
T = 25.8E7E32Z*T
SN12 = (((((5.E4E-1E*T - 1.3747F-13)*T + 3.11917E-11)*T
1 - 6.434E09E-9)*T + 1.2L7F2E-6)*T - 2.57626E-4)*T
2 - 0.0826424E + 0.019849E4*ALCG(T))*T + 1.0)*0.5
RETURN
110 CONTINUE
IF (T .LT. 4.0)  GC TC 120
X = (T - 2.5)/1.5
SN12 = ((((((5.0E5E1E-3 - 1.4146EE-3*X)*X - 2.98085F-3)*X
1 + 9.48136E-3)*X - 0.022794E7)*X + 0.0523088)*X
2 - 0.1045084)*X + 0.18E2743)*0.5
RETURN
120 CONTINUE
IF (T .LT. 10.0)  GC TC 130
X = (T - 7.0)/3.0
SN12 = (((((5.08470E-4 - 1.00181E-4*X)*X - 5.27E94E-4)*X
1 + 2.24274E-3)*X - 6.5E42C8E-2)*X + 0.01290804)*X
2 - 0.02622E7)*X + 0.0E5522E7)*0.5
RETURN
130 CONTINUE
IF (T.GT.100.0)  GC TC 140
Y = ALOG(T)
X = 5.210340371976/Y - 3.0
SN12 = (((((7.836E-2*X - 1.5181E-3)*X + 9.0128E-4)*X
1 - 6.49E4E-4)*X - 1.0E32E-3)*X + 0.01493E4)*X
2 + 0.4372273)*0.5/T
RETURN
140 CONTINUE
Y = ALOG(T)
X = 5.210340371976/Y - 1.0
SN12 = (((((5.074EE-3 - 4.681E-4*X)*X - 1.5848E-3)*X
1 + 0.014752F)*X + 0.4079074)*0.5/T
RETURN
200 CONTINUE
IF (T .LT. 1.0)  GC TC 210
T = 25.0729185*T
SN12 = (((((2.71E-1E*T - 6.9372E-14)*T + 1.58314E-11)*T
1 - 3.39873E-9)*T + 7.01818E-7)*T - 1.01594E-4)*T
2 - 0.065E453 + 0.01494284*ALCG(T))*T + 1.0)*0.5
RETURN
210 CONTINUE
IF (T .LT. 4.0)  GC TC 220
X = (T - 2.5)/1.5
SN12 = (((((2.32E-3 - 1.123E-3*X)*X - 3.0489E-3)*X
1 + 7.934E-3)*X - 8.021328E1)*X + 0.0521E86)*X
2 - 0.1212E60)*X + 0.25089021)*0.5
RETURN
220 CONTINUE
IF (T .LT. 10.0)  GC TC 230
X = (T - 7.0)/3.0
SN12 = (((((4.727E-4*X - 1.2459E-3)*X + 2.52E3E-3)*X
1 - 6.7323E-3)*X + 0.017476E1)*X - 0.0421443)*X
2 + 0.095E415)*0.5
RETURN
230 CONTINUE
IF (T .LT. 100.0)  GC TC 240
Y = ALOG(T)
X = 5.210340371976/Y - 3.0
SN12 = (((((9.1E-5*X - 1.809EE-3)*X - 1.8059F-3)*X
1 + 0.0263129)*X + 0.0392594)*0.5/T
RETURN
240 CONTINUE
Y = ALOG(T)
X = 5.210340371976/Y - 1.0
SN12 = (((((6.444E-4*X - 2.4E3E-4)*X - 5.099E-4)*X
1 + 0.02161E0)*X + 0.05910521)*0.5/T
RETURN
200 CONTINUE
SN12 = 0.0
RETURN
END

```

```

FUNCTION TAUMAX (KL,KE,FM,PV,T,FLAM,S,DEN)
SIG = 1.43870*PV/T
TAUMAX = 5.3E13E-1E*SGRT(SM/T)*FLAM*DEN*S*SJMAX(KL,KE,SIG)
RETURN
END

C FUNCTION SJMAX (KL,KE,SIG)
S,MAX COMPUTES THE RELATIVE STRENGTH OF THE STRONGEST LINE
IF KB = 1, SIGMA-SIGMA TRANSITION
IF KB = 2, SIGMA-PI TRANSITION
GO TC (110,120,130), KE
110 CONTINUE
TJM = 0.5*(SGRT(C,2E+2,0/SIG) - 1.5)
A = C.5*(TJM + TJM + 1.0)
GO TC 200
120 CONTINUE
TJM = 0.5*(SGRT(2.0/SIG) - 1.0)
E = TJM + 1.0
GO TC 200
130 CONTINUE
SJMAX = 0.0
RETURN
200 CONTINUE
SJM = A*EXP(-SIG*TJM*(TJM + 1.0))
SJMAX = SJM/CJFAPT(KL,SIG)
RETURN
END

C FUNCTION CJFAPT (K,SIG)
CJFAPT COMPUTES THE ROTATIONAL PARTITION FUNCTION FOR A
RIGID ROTATOR, SIG .LE. 0.2.
K = 1, ALL LEVELS FCUPULATEC
K = 2, EVEN LEVELS FCUPULATEC
K = 3, ODD LEVELS FCULLATEC
IF (SIG .GT. 0.2) STCF
X = SIG
IF (K = 2) 110,120,130
COEFFICIENTS ARE FROM RUN CALTE-EV
110 CONTINUE
CJFAPT = (((((0.0038E5C0232E57*SIG + 0.0125718235471)*X
1 + 0.06E6E75E1E4520)*X + 0.232332133372)*X + 1.00000000017)/SIG
RETURN
120 CONTINUE
CJFAPT = (((((62.5879214103*SIG - 29.5697177693)*SIG
1 + 6.49E52E700065*SIG - 0.202E72235E08)*SIG
2 + 0.02E737E193546)*SIG + 0.0328350E09506)*SIG
3 + 0.1E6E670220770)*SIG + 0.49999999456)/SIG
RETURN
130 CONTINUE
CJFAPT = (((((29.57211787E - 62.5E7459E502*SIG)*SIG
1 - 6.49E57E373665)*SIG + 0.5E5750942701)*SIG
2 - 0.01E6E7551E7E)*SIG + 0.0328315E8E0E9)*SIG
3 + 0.1E6E3112387)*SIG + 0.5E000000E44)/SIG
RETURN
END

FUNCTION SIMP (A,H,K)
DIMENSION A(3)
IF (K = 2) 100,110,120
100 CONTINUE
SIMP = (8.0*A(2) + 5.0*A(1) - A(3))*H/12.0
RETURN
110 CONTINUE
SIMP = (8.0*A(2) + 5.0*A(3) - A(1))*H/12.0
RETURN
120 CONTINUE
SIMP = (4.0*A(2) + A(1) + A(3))*H/3.0
RETURN
END

```

```

FUNCTION SM1 (K, TAU)
I = ABS(TAU)
IF (T .EG. C.0) GO TC 300
IF (K .GE. 2) GO TC 200
IF (T .GT. 4.0) GO TC 110
X = (T - 2.0)/4.0
SM1 = (((((2.35450E-03)*X - 8.90535E-03)*X + 2.56223E-(2)*X
1 - 7.07E77E-02)*X + 1.63735E-01)*X - 3.01749E-01)*X + 0.426845
RETURN
110 CONTINUE
IF (T .GT. 10.0) GO TC 120
X = (T - 7.0)/3.0
SM1 = (((((5.92257E-04)*X - 1.79043E-03)*X + 4.05109E-03)*X
1 - 1.04749E-02)*X + 2.59109E-02)*X - 6.01702E-02)*X + 0.134123
RETURN
120 CONTINUE
IF (T .GT. 100.0) GO TC 130
Y = ALOG(T)
X = 6.210240371976/Y - 3.0
SM1 = (((((-6.24516E-04)*X - 6.79378E-04)*X + 2.98590E-03)*X
1 - 1.28150E-03)*X - 6.31784E-03)*X + 3.78904E-02)*X
2 + 8.90441E-01)/T
RETURN
130 CONTINUE
Y = ALOG(T)
X = 6.210240371976/Y - 1.0
SM1 = (((((6.12038E-04)*X - 2.93434E-04)*X - 4.36163E-04)*X
1 + 1.30037E-02)*X - 3.00082E-04)*X + 3.23441E-02)*X
2 + 8.17407E-01)/T
RETURN
200 CONTINUE
IF (T .GT. 4.0) GO TC 210
X = (T - 2.0)*0.5
SM1 = (((((1.39071E-03)*X - 5.25004E-03)*X + 1.58249E-02)*X
1 - 4.74039E-02)*X + 1.26175E-01)*X - 2.06333E-01)*X + 0.517573
RETURN
210 CONTINUE
IF (T .GT. 10.0) GO TC 220
X = (T - 7.0)/3.0
SM1 = (((((3.76420E-04)*X - 1.25384E-03)*X + 3.411E8E-03)*X
1 - 1.05380E-02)*X + 3.07476E-02)*X - 8.14705E-02)*X + 0.1941308
RETURN
220 CONTINUE
IF (T .GT. 100.0) GO TC 230
Y = ALOG(T)
X = 6.210240371976/Y - 2.0
SM1 = ((((-7.53415E-03)*X - 8.87124E-03)*X + 6.62199E-02)*X
1 + 1.304245)/T
RETURN
230 CONTINUE
Y = ALOG(T)
X = 6.210240371976/Y - 1.0
SM1 = (((((1.09067E-03) - 5.82E10E-04)*X)*X + 1.59569E-03)*X
1 + 1.21174E-03)*X + 4.76165E-02)*X + 1.18445)/T
RETURN
300 CONTINUE
SM1 = 1.0
RETURN
END

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DC 790 M=M1,MSET
IF(M.GT.C1) GC TO 770
C2=M+E9+E371
GC TO 780
770 C2=150+(M-91)*5+E371
780 ANG=ASIN(SIN(PHI)*[1/C2])
CS=SQRT([C1+C1+D2*D2-2]*[C1*D2*(CS(PHI-ANG))]
DELS=(DS-RANG)*1.E+15
RANG=CS
DELTAU = DELS*(0.E*(TAL(M,J)+TAL(M-1,J)))
STAU(L) = STAU(L) + DELTAU
SL1E = SL1(KTAU,STAU(L))
RADA = RAD(M-1)/TAU(M-1,J)
FADE = FAC(M)/TAL(M,J)
SFAC = (RACE - RACA)/DELTAU
SFAD(L) = SRAD(L) + SFAC*(SL1A - SL1B)
SL1A = SL1E
SAVTAU=STAU(L)
790 CCNTINUE
SFAD(L) = SRAD(L) + RACE*SLC(KTAU,SAVTAU)
E10 CCNTINUE
SAVTAL=0.0
STAU(7)=0.0
SL1A=0.0
VRANG=0.0
SRAD(8)=0.0
MSET=IALT
DC 815 M=M+MSET
L=MSET-M+K
IF(L.GT.C1) GC TO E11
E2=L+E9+E371
GC TO E12
E11 C2=150+(L-C1)*5+E371
E12 DS=SQRT([C2*C2-D1*[1])

IF(L.LT.MSET) GO TO E13
RANG=CS
GC TO E13
E13 DELS=(RANG-CS)*1.E+15
VRANG = VRANG + DELS
RANG=CS
SAVTAU = STAU(7)
DELTAU = DELS*(0.E*(TAU(L,J)+TAL(L+1,J)))
STAU(7) = STAU(7) + DELTAU
SL1B = SL1(KTAU,STAU(7))
RADOB = RAD(L)/TAL(L,J)
RADA = RAD(L+1)/TAU(L+1,J)
SFAC = (RACE - RACA)/DELTAU
SFAD(8) = SRAD(8) + SFAC*(SL1A - SL1B)
SL1A = SL1E
E15 CCNTINUE
M1=K+1
RANG=0.0
MSET=IALT
DC 816 M=M1,MSET
IF(M.GT.C1) GC TO E16
E2=M+E9+E371
GC TO E16
E16 C2=150+(M-C1)*5+E371
E17 CS=SQRT([C2*C2-D1*[1])
DELS=(DS-RANG)*1.E+15
VRANG = VRANG + DELS
RANG=CS
SAVTAU = STAU(7)
DELTAU = DELS*(0.E*(TAU(M,J)+TAL(M-1,J)))
STAU(7) = STAU(7) + DELTAU
SL1B = SL1(KTAU,STAU(7))
RADOB = RAD(M)/TAL(M,J)
RADA = RAD(M-1)/TAU(M-1,J)
SFAC = (RACE - RACA)/DELTAU
SFAD(8) = SRAD(8) + SFAC*(SL1A - SL1B)
SL1A = SL1E
E18 CCNTINUE
SFAC(8) = SRAD(8) + RACE*SLC(KTAU,STAU(7))
WRITE(3,4444) HT,SRAC(8),SRAC(1),SRAD(7)
4444 FFORMAT(FE8.1,P7.4)
WRITE(6,7409) HT,SRAC(8),HT,(SRAC(M), M= 1,7)
E30 CCNTINUE
E40 CCNTINUE
STOP 77
END

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FUNCTION SLO (K, TAU)
T = ABS(TAU)
IF (K .EQ. 2) GC TC 200
IF (T .GT. 4.0) GC TC 110
X = (T - 2.0)*0.5
1 SLO = (((((0.0014110*x - 0.011496)*x + 0.0068378)*x +
- 0.0240060)*x + 0.0792553)*x - 0.2401023)*x + 0.6569377)*T
RETURN
110 CONTINUE
IF (T .LT. 10.0) GC TC 120
X = (T - 7.0)/3.0
1 SLO = (((((0.001918*x - 0.007254)*x + 0.0023078)*x -
- 0.0080248)*x + 0.0274865)*x - 0.0942082)*x + 0.3539295)*T
RETURN
120 CONTINUE
Y = ALOC(T)
IF (T .LT. 100.0) GC TC 130
X = S.210340371976/Y - 3.0
1 SLO = (((((0.001490*x - 0.0001530)*x - 0.001490*x +
+ 0.0013443)*x - 0.0003060)*x + 0.0837063)*x + 1.1415564)*Y
RETURN
130 CONTINUE
X = S.210340371976/Y - 1.0
1 SLO = (((((0.0017142*x - 0.00222*x) + 0.005211)*x -
- 0.004886)*x - 0.006477)*x + 0.007743)*x - 0.015900)*X
2 + 0.122528)*x + 0.040111)*Y
RETURN
200 CONTINUE
IF (T .GT. 4.0) GC TC 210
X = (T - 2.0)*0.5
1 SLO = (((((0.002431*x - 0.010901)*x + 0.0041830)*x -
- 0.0160764)*x + 0.0581766)*x - 0.2013270)*x + 0.7188944)*T
RETURN
210 CONTINUE
IF (T .LT. 10.0) GC TC 220
X = (T - 7.0)/3.0
1 SLO = (((((0.0001169*x - 0.0004699)*x + 0.0016637)*x -
- 0.0005345)*x + 0.0255562)*x - 0.1003592)*x + 0.4282944)*T
RETURN
220 CONTINUE
Y = ALOC(T)
X = S.0017018598805/Y - 1.0
1 SLO = (((((0.036251*x - 0.035886*x)*x + 0.062211)*x -
- 0.056845)*x - 0.048818)*x + 0.045072)*x + 0.005249)*X
2 - 0.000887)*x - 0.062056)*x + 0.162331)*x + 1.399785)*Y
END

FUNCTION SL1 (K, TAU)
IF (K .EQ. 1) COMPUTE FUNCTION VALUE FOR PARALLEL BAND.
IF (K .EQ. 2) COMPUTE FUNCTION VALUE FOR FERFENDICULAR BAND.

T = ABS(TAU)
IF (K .GE. 2) GC TC 200
IF (T .GT. 1.0) GC TC 110
X = (T - 0.5)*2.0
1 SL1 = (((((1.9E-5*x - 1.4E25E-4)*x + 0.00276149)*x -
- 0.036251)*x + 0.4135671)*T*T
RETURN
110 CONTINUE
IF (T .GT. 4.0) GC TC 120
X = (T - 2.5)/1.0
1 SL1 = (((((3.497E-4 - 5.4CE-5*x)*x - 2.0328E-3)*x +
+ 0.0114067)*x - 0.0613564)*x + 0.3531337)*T*T
RETURN
120 CONTINUE
IF (T .GT. 10.0) GC TC 130
X = (T - 7.0)/3.0
1 SL1 = (((((7.771E-4 - 1.878E-4*x)*x - 2.9392E-3)*x +
+ 0.0120034)*x - 0.051138)*x + 0.2354310)*T*T
RETURN
130 CONTINUE
IF (T .GT. 100.0) GC TC 140
Y = ALOC(T)
X = S.210340371976/Y - 3.0
1 SL1 = (((((F.67E-5*x - E.7EFE-4)*x + 0.0021934)*x -
+ 0.0011841)*x + 0.058282)*x + 0.8538766)*T*Y
RETURN
140 CONTINUE
IF (T .GT. 1000.0) GC TC 150
Y = ALOC(T)

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X = 27.63102111553/Y - 5.0
SL1 = (((1.636E-4*X - 1.5E04E-3)*X + 1.6892E-3)*X
1 + 0.861629E)*T*Y
150 RETURN
CONTINUE
Y = ALOC(T)
X = 13.6155105575E/Y - 1.0
SL1 = (((((2.5E-4*X + 2.709E-3)*X - 7.600E-3)*X
1 + 6.917E-3)*X + 2.37E-4)*X - 1.2795E-2)*X
2 + 0.030224)*X + 0.838671)*T*Y
RETURN
200 CONTINUE
IF (T .GT. 1.0) GC TC 210
X = (T - 0.5)*2.0
SL1 = (((((2.5E-4*X - 1.2125E-4)*X + 1.87580E-3)*X
1 - 0.02632252)*X + 0.4667332)*T*T
RETURN
210 CONTINUE
IF (T .LT. 2.5/1.5) GC TC 220
X = (T - 2.5/1.5)
SL1 = (((((2.14E-4 - 3.26E-5)*X - 1.3290E-3)*X
1 + 8.3419E-3)*X - 0.0530254)*X + 0.2801694)*T*T
RETURN
220 CONTINUE
IF (T .GT. 10.0) GC TC 230
X = (T - 7.0)/3.0
SL1 = (((((5.230E-4 - 1.1E2E-4*X)*X - 2.2120E-3)*X
1 + 0.0102E27)*X - 0.04542E21)*Y + 0.2718101)*T*T
RETURN
230 CONTINUE
IF (T .GT. 100.0) GC TC 240
Y = ALOC(T)
X = 210740371976/Y - 3.0
SL1 = (((((1.06PE-4*X - 1.2224E-3)*X + 3.8304E-3)*X
1 + 8.1841E-3)*X - 6.6E22E-2)*X + 1.057781)*T*Y
RETURN
240 CONTINUE
IF (T .GT. 1000.0) GC TC 250
Y = ALOC(T)
X = 27.63102111553/Y - 5.0
SL1 = (((2.668E-4*X - 2.2012E-3)*X
1 - 0.0181E06)*X + 1.1468715)*T*Y
RETURN
250 CONTINUE
Y = ALOC(T)
X = 13.6155105575E/Y - 1.0
SL1 = (((((4.9E9F-4*X - 2.2191F-3)*X + 3.6173E-3)*X
1 - 2.8417E-3)*X + 3.34E2E-3)*X - 0.0192525)*X
2 + 3.6151E-3)*X + 1.1747922)*T*Y
RETURN
END

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PROGRAM SFCTEA1 (CL1FL1,TAFEL,TAPF1,TAPF2,TAFE5)
CIMENST(N,LSC(2E),LCC(2E),FCM(2E),RV(25),LPC(25),LBL(25),STR(25)
CIMENSI(N,CCN(20,10),TEMP(20),FRDE(20))
CIMENSI(N,AVTEMP(40,25),EVTEMP(40,25))
CIMENSI(N,CCLCNT(40,25),HRCNT(40,25))
CIMENSI(N,FACE(50),RACE(50),RADC(50))
COMMCN /ISFCCM/ ISF
READ (1,12) TEMP
CC 500 NCASES = 1, 2
REWIND E
READ (1,E) NSPTOT
DC 400 NSF = 1, NSFTOT
READ (1,E) ISP
READ (1,41) NLEVEL,NEANC,RM
CC 210 I = 1, NLEVEL
READ (1,41) LSC(I),LCC(I),FCM(I),RV(I)
10 CCNTINUE
CC 220 I = 1, NPARC
READ (1,41) LPC(I),LBL(I),STR(I)
220 CCNTINUE
READ (1,E) NALT,NLEV
CC 221 N2 = 1, NALT
READ (1,42) CCN(NZ,NL), NL = 1, NLEVEL
221 CCNTINUE
COMPUTE MEAN TEMPERATURES AND COLUMN COUNTS
K = NLEV
CC 120 NL = 1, NLEVEL
CC 112 I = 1, NALT
FRCF(I) = CCN(I,NL)*TFMF(I)
112 CCNTINUE
TF = 0.0
TC = 0.0
NA = NALT
CC 120 I = 1, K
TF = TF + SIMP(PFCE(NA - 2),1.0E+5,2)
1 + SIMP(FRCF(NA - 3),1.0E+5,3) + SIMP(FCOC(NA - 5),1.0E+5,3)
1 TC = TC + SIMP(CCN(NA - 2,NL),1.0E+5,2)
1 + STMF(CCN(NA - 3,NL),1.0E+5,2) + SIMP(CON(NA - 5,NL),1.0E+5,2)
COLCNT(K - I + 1,NL) = TC
AVTEMP(K - I + 1,NL) = TF/TC
NA = NA - E
FORCANT(I,NL) = 0.0
120 CCNTINUE
REASE = 6.431E+8
CC 120 I = 1, K
TOP = 0.0
TCH = 0.0
NA = NALT
RR2 = REASE*RBASE
ETOP = 6.431E+8 * NALT*1.0E+5
SVH = SGRT(RTOP*RTCF - RE2)
KK = K - I + 1
CC 125 J = 1, KK
RTOP = RTOP - 1.0E+5
STH = SGRT(RTOP*RTCF - RE2)
HF = SVH - STH
SVH = STH
RTOP = RTOP - 1.0E+5
STH = SGRT(RTOP*RTCF - RE2)
F2 = SVH - STH
SVH = STH
RTOP = RTOP - 1.0E+5
STH = SGRT(RTOP*RTCF - RE2)
F1 = SVH - STH
SVH = STH
TFH = TFH + QINT(FRCC(NA - 2),H4,H5,2)
1 + QINT(FRCC(NA - 3),H3,H4,2) + QINT(FRCC(NA - 5),H1,H2,3)
1 TCH = TCF + QINT(CCN(NA - 2,NL),H4,H5,2)
1 + QINT(CCN(NA - 3,NL),H3,H4,3) + QINT(CCN(NA - 5,NL),H1,H2,3)
NA = NA - 5
125 CCNTINUE
FORCANT(I,NL) = TCH
EVTEMP(I,NL) = TFH/TCH

```

```

128  ERASE = REASE + 5.0E+5
129  CCNTINUE
130  CC 400 NE = 1, NBAND
131  CC 230 I = 1, NRLEV
132  READ (1,43) HT,PAEC(I),PACA(I),PADB(I)
133  CCNTINUE
134  WRITE(4) (LSC(I),I=1,25)
135  WRITE(4) ((FCRCNT(I,J),J=1,25),I=1,40)
136  WRITE(4) ((EVTEMP(I,J),J=1,25),I=1,40)
137  400  CCNTINUE
138  500  CCNTINUE
139  STOP
140  6   FCRMAT (E16)
141  P   FCRMAT (1FEE11.3)
142  14  FCRMAT (1X)
143  12  FCRMAT (1FFEE12.4)
144  18  FCRMAT (1F1)
145  41  FCFMAT (2IE,5E12.4)
146  42  FCRMAT (8F10.4)
147  43  FCRMAT (FE.1,3E17.4)
END

```

```

FUNCTION SIMP (A,F,K)
DIMENSION A(3)
IF (K = 2) 100,110,120
100 CCNTINUE
SIMP = (2.0*A(2) + 5.0*A(1) - A(3))*H/12.0
RETURN
110 CCNTINUE
SIMP = (2.0*A(2) + 5.0*A(3) - A(1))*H/12.0
RETURN
120 CCNTINUE
SIMP = (4.0*A(2) + A(1) + A(3))*H/3.0
RETURN
END

```

```

FUNCTION GINT (A,F1,F2,K)
DIMENSION A(3)
HS = H1 + H2
IF (K = 2) 100,110,120
100 CCNTINUE
GINT = ((A(2)*HS - A(1)*H2 - A(3)*H1)*H1/(6.0*H2*HS)
1 + (A(1) + A(2))*C.5)*F1
RETURN
110 CCNTINUE
GINT = ((A(2)*HS - A(1)*H2 - A(3)*H1)*H2/(6.0*H1*HS)
1 + (A(2) + A(3))*C.5)*F2
RETURN
120 CCNTINUE
GINT = (A(2)*HS - A(1)*H2 - A(3)*H1)*(H1**3 + H2**3)
1 /(6.0*HS*F1*H2) + (A(1)*H1 + A(2)*HS + A(3)*H2)*C.5
RETURN
END

```

```

PROGRAM SFCTRA2 (CUTPUT,TAFE4,TAPE1,TAPE2,TAFE5)
DIMENSION LEC(25), LSC(25)
DIMENSION CCN(10), TEMP(1201)
DIMENSION FCRCNT(40, 25), EVTEMP(40, 25)
DIMENSION FACAS(40), RACE(50), RADC(50)
COMMON /ISFCOM/ ISP
COMMON /WVNLIM(2500), ENLS(2500), STR(2500), HRAC(2500)
COMMON /SFEC/ 1801
GO 600 J=1,14
REWINE 1
REWINE 4
    RFAC(1,12) TEMP
DC 500 NCASES = 1, 2
REWINE 5
CC 100 I = 1, 1801
SPEC(I) = 0.0
100 CCNTINUE
    READ(1,6) NSPTCT
DC 400 NSF = 1, NSFTCT
    READ(1,6) ISP
    READ(1,41) NLEVEL,NEANC,RM
    READ(4)(LSC(I),I=1,25)
    READ(4)(FCRCNT(I,JKK),JKK=1,25),I=1,40)
    READ(4)(EVTEMP(I,JKK),JKK=1,25),I=1,40)
CC 210 I = 1, NLEVEL
    READ(1,41) LSCC,LCC,RCM,EV
210 CCNTINUE
    DC 220 I = 1, NEANC
    READ(1,41) LBC(I),LBU,STR
220 CCNTINUE
    READ(1,6) NALT,NRLEV
DC 221 NZ = 1, NALT
    READ(1,42) (CCN(NL),NL=1,NLEVEL)
    CONTINUE
    IF (ISF .NE. 3) GO TC 225
    CNTHR = 2.0*HOPCNT(J,1)
    CALL ROTATE(5,LINES)
C    RCTATF WILL READ FCRCM TAPE5
    CALL WATSTR(RADC(J),RM,EVTEMP(J,1),CNTHR,1,LINES)
    CALL FILTER(SPEC,RFAC,WVNLIM,LINES)
225 CCNTINUE
    DC 400 NE=1,NEANC
    DC 230 I=1,NRLEV
    READ(1,43) FT,RADC(I),RACA(I),RACE(I)
230 CCNTINUE
    NEC=LEC(NE)
    NI=MOC(NEC/100,100)
    NF=MOC(NEC,100)
    IF ((NB .GT. 10) .AND. (ISF .EG. 3)) GO TC 400
        CALL ROTATE(5,LINES)
        CALL WATSTR(RADC(J),RM,BVTEMP(J,NF),2.0*HOPCNT(J,NF),2,LINES)
400 CCNTINUE
    WRITE(2,16)
    WRITE(2,8) SPEC
500 CCNTINUE

600 CCNTINUE
    STOP
    6 FORMAT(E1E)
    8 FORMAT(1FF11.3)
    14 FORMAT(1X)
    12 FORMAT(1FEE12.4)
    1E FORMAT(1F1)
    41 FORMAT(2IE,5E12.4)
    42 FORMAT(PE1C.4)
    43 FORMAT(FE.1,2E17.4)
END

SUBROUTINE WATSTR (FAC,RM,T,CEN,N,LINES)
COMMON /ISFCOM/ ISP
COMMON /WVNLIM(2500), ENLS(2500), STR(2500), HRAC(2500)
TOFAC = (25E.0/T)**1.5
IF (ISF .EG. 2) TGFAC = 25E.0/T
IF (ISF .EG. 4) TGFAC = 25E.0/T
GO TC (100,200,300), R
100 CCNTINUE
    FADFFC = 1.5E05E-20*CEN*TGFAC
    FAUFAO = 4.8815E-18*SGT(PM/T)*TGFAC*CEN
    DC 210 I = 1, LINES

```

```

HHOLAM = 1.0E+4/HVNLIM(I)
TAU = TALFAC*STR(I)*EXP(-1.43879*ENLS(I)/T)
1 * (1.0 - EXP(-14387.9/(HHOLAM*T)))
ROTEA = 2.0*STR(I)/HHCLAM*#2
1 FRAD(I) = RACFAC*FCTEA*SFUN(TAU)
1 * EXP(-1.43879*(1.0E+4/HHCLAM + ENLS(I))/T)
110 CCNTINUE
RETURN
200 CCNTINUE
TAUFAC = 4.2815E-18*SGET(RM/T)*CEN
TEFAC = (T - 296.0)/(T+296.0)*1.43879
SUM = 0.0
CC 210 I = 1, LINES
FSTR = TGFACT*STR(I)*EXP(TEFAC*ENLS(I))
HHOLAM = 1.0E+4/HVNLIM(I)
TAU = TALFAC*HHCLAM*HSTR
1 FRAC(I) = TAL*SFLN(TAL)*EXP(-14387.9/(HHOLAM*T))
1 /HHCLAM*#4
1 SUM = SLM + HRAD(I)
210 CCNTINUE
RADFAC = FRAC/SUM
CC 220 I = 1, LINES
FFAC(I) = HRAD(I)*RACFAC
220 CCNTINUE
RETURN
300 CCNTINUE
RETURN
END

FUNCTION SFUN (TAL)
T = ABS(TAL)
IF (T .LT. 4.0) GO TO 110
X = (T - 2.0)*0.5
SFUN = (((((0.0010E1*X - 0.004288)*X + 0.013686)*X - 0.042491)
1 * X + 0.0114932)*X - 0.2E70E7)*X + 0.556485
RETURN
110 CCNTINUE
IF (T .GT. 10.0) GO TO 120
X = (T - 7.0)/3.0
SFUN = (((((0.00347*X - 0.001167)*X + 0.003148)*X - 0.005545)
1 * X + 0.028248)*X - 0.081972)*X + 0.247900
120 RETURN
Y = ALOC(T)
X = 1.0/Y
SFUN = (((((0.0E787E4*X - 0.2EPE58)*X + 0.2645E8)*X
1 - 0.2691E71)*X + 0.32E2E8)*X + 1.128379)*SQR(Y)/T
RETURN
END

C SUBRCUT THE FILTER (SPEC, RAD, HVNUM, N)
FIVE WAVE NUMBER RESOLUTION, TWO WAVE NUMBER INTERVAL
DIMENSION SPEC(1E01), RAD(4E01), HVNUM(4E01)
CCMMCN /ISFCOM/ ISF
CC 150 I = 1, N
RT = HVNLIM(I)
RADT = RAD(I)
JS = (RT - 405.0)*0.5 + 2.0
JF = (RT - 305.0)*0.5 + 1.0
IF (JS .LT. 1) JS = 1
IF (JF .GT. 1801) JF = 1801
IF (JS .GT. JF) GO TO 150
CC 120 J = JS, JF
ENUM = (J - 1)*?^P + 4[0.0
SPEC(J) = SPEC(J) + (5.0 - AES(HVNUM - RT))*RADT*0.04
120 CCNTINUE
150 CCNTINUE
150 RETURN
END

SUBRCUT INF.RCTATE (NIN,LINES)
CCMMCN /ISFCOM/ ISF
CCMMCN HVNLIM(2500), ENLS(2500), STR(2500), HFAC(2500)
READ (NIN) LINES
1 IF (ECF(NIN)) 121, 100
190 READ (NIN)(HVNUM(I)), STR(I), ENLS(I), I=1, LINES
1 IF (ECF(NIN)) 121, 115
119 CCNTINUE
121 RETURN
END

```

Appendix B

```

C      PROGRAM PLCTK(INPUT,OUTPUT=/E000,TAPE1=INPUT,TAPE2=CLTPLT,TAPE3)
LAST CHANGE FEB 15, 1979

      DIMENSION K(1E)
      DATA COL /"6"/
1      FFWINC 3
      READ(3,500) TYPE
500    FCFMAT(A1,A7,2I8,ZF18.11)
      IF(TYPE.EQ.COL) GO TO 10
      REWIND 3

      WRITE(7,1000)
1000  FCFMAT(* ENTER DELT,START,STCF,AMP,OFFSET,STAR10*)
      READ(1,* ) DELT,WIND,XEND,AMF,CFF,START
      GO TO 15
10     REWIND 3
      READ(3,500) TYPE,TITLE,NUM,IREV,AE,DELT,START
      WRITE(2,550)
550    FCFMAT(* ENTER START,STCF,AMF,OFFSET*)
      READ(1,* ) WIND,XEND,AMF,CFF

15     REWIND 2
      XDELT=PC0./(XEND-WIND)
      XDELT=DELT*XDELT
      WRITE(2,560) TITLE
560    FCFMAT(/,15X,A7)
      WPI TE(2,20000)WIND,XEND
20000  FORMAT(3X,INTERVAL BETWEEN *,F8.3,* AND *,F8.3)
      WPI TE(2,30000) CFF,AMF
30000  FCFMAT(3X,*AMPLITUDE RANGE(*,F7.0,* TO *,F7.0,*))

      REWIND 2
      CALL FLOTS
      CALL FRAME

      IY2=8
      IX2=10
      X2=10
      XN=START

      DO 100 I=1,10000
      READ(3,40000)(K(J),J=1,1E)
+0000  FORMAT(16I5)
      IF(ECF(3)) 900,80,P0
      X=XN
      XN=XN+16*DELT
      IF(XN.LT.WIND)GOTC 100
      EN 80 J=1,16
      XJ=X+FLOAT(J-1)*DELT
      IF(XJ.LT.WIND)GOTC 80

      IX=IX2
      IY=IY2
      X2=X2+XDELT
      IX2=IX2
      IF(IX2.GT.900)GOTC 800
      IY2=50J.+((K(J)-CFF)/(AMP-CFF)+R.5
      IF(IY2.LE.1) IY2=1
      CALL LINE(IX,IY,IY2,IY2)
      CONTINUE
      CONTINUE

800    REWIND 3
      CALL XMTT
      READ(1,* )JUNK
      REWIND 2
      GOTC 1
      STOP 7
      END

```

THIS ROUTINE INITIALIZES THE PLOT BUFFER
AND THE INDEXES INTO THE PLOT BUFFER.

SUBROUTINE PLCTS

```
COMMON/IXY/IT,ICX,ICLY,ICHY,TOLX,IOHX  
COMMON/BUFFER/LB,LW,IB(500)
```

```
LE=1  
IT=0  
LW=3E  
IE(1)=00070000000000000000000000000000  
RETURN  
END
```

THIS SUBROUTINE PACKS A 12-BIT
PLOT INSTRUCTION FOR THE TEKTRONIX INTO
THE BUFFER (IB). IT IS FORMATTED TO THE
PECULIAR BIT STREAM REQUIREMENTS FOR TEMPO.

```
SUBROUTINE SEND(I)  
COMMON/BUFFER/LB,LW,IB(500)  
IF(LW.GE.0)GOTO 2  
LW=4B  
LB=LB+1  
2 IB(LB)=OR(AND(IE(LB),SHIFT(MASK(4E),LW)),SHIFT(I,LW))  
LW=LW-12  
1 FORMAT(3X,I4,3X,I4)  
RETURN  
END
```

```
SUBROUTINE XMTT  
COMMON/IXY/IT,TCX,ICLY,ICHY,IOLX,IOHX  
COMMON/BUFFER/LB,LW,IB(500)
```

```
CALL SEND(ICHY)  
CALL SEND(ICLY)  
CALL SEND(IOHX)  
CALL SEND(IOLX)  
DO 1 K=1,9  
1 IF(LW.EQ.0)GOTO 2  
CALL SEND(0)  
2 BUFFER OUT(2,0) (IB(1),IB(LB))  
IF(UNIT(2)) 3,3,3  
3 CALL FLCTS  
RETURN  
END
```

SUBROUTINE PACF

```
CALL SEND(4033E)  
CALL SEND(4014E)  
CALL SEND(4007E)  
CALL SEND(0)  
CALL XMIT  
RETURN  
END
```

C

SUPPOLTINE COPY

```
CALL SEND(4033B)
CALL SEND(4027E)
CALL XMIT
RETURN
END
```

X

```
SUBROUTINE FRAME
CALL LINE(10,500,900,509)
CALL LINE(10,768,900,768)
CALL LINE(10,8,900,8)
CALL LINE(9,768,9,9)
CALL LINE(898,768,898,9)
CALL LINE(10,768,900,768)
CALL LINE(10,768,800,7F7)
CALL LINE(10,10,500,10)
CALL LINE(10,5,900,5)
CALL LINE(10,768,10,9)
CALL LINE(11,768,11,6)
CALL LINE(900,768,900,9)
CALL LINE(898,768,899,9)
RETURN
END
```

C

```
SUPEROLINE LINE(J,I,JJ,IT)
COMM CN/EUFFER/LB
CALL SEND(4035B)
```

```
K=OR(AND SHIFT(I,-5),-MASK(55)),4040B
CALL SEND(K)
K=OR(AND(I,-MASK(55)),4140E)
CALL SEND(K)
K=OR(AND SHIFT(J,-5),-MASK(55)),4040B
CALL SEND(K)
K=OR(AND(J,-MASK(55)),4100B)
CALL SEND(K)
K=OR(AND SHIFT(II,-5),-MASK(55)),4040B
CALL SEND(K)
K=OR(AND(II,-MASK(55)),4140E)
CALL SEND(K)
K=OR(AND SHIFT(JJ,-5),-MASK(55)),4040B
CALL SEND(K)
K=OR(AND(JJ,-MASK(55)),4100E)
CALL SEND(K)

IF(LB.GT.480)CALL XMT
FETLFN
END
```

Appendix C

```

*****  

*****  

PLOTR EQU 0F5H  

PLTCS EQU 0F7H  

MDMCS EQU 0EFH  

MDMDT EQU 0EEH  

TXRDY EQU 01  

RXRDY EQU 02  

BMOD0 EQU 80H  

EOCHR EQU 83H  

NDCHR EQU 0FFH  

LFTPEN EQU 40H  

DRPEN EQU 80H  

LWAIT EQU 50 : TIME DELAY FOR PEN UP/DOWN  

SWAIT EQU 255  

OVRIT EQU 10H  

DICMP EQU 10010001B  

NXORG EQU 0600 :  

*****  

ORG 0800H  

*****  

IN MDMDT : CLEARS MODEM  

MVI A, "R" : CYBER RHN COMMAND  

CALL XMIT  

MVI A, "N"  

CALL XMIT  

MVI A, "H"  

CALL XMIT  

CALL SNDCR : Xmits 'RETURN'  

CALL HEXIN : IGNORE CR,LF:WAIT FOR PROMPT  

MVI A, BMOD0 : INIT PLOTTER PORT  

SET55 EQU $  

OUT PLTCS  

LXI H, BUFER : RESET PNTER  

SHLD PTRER  

LXI H, 0 : CLR CURRX, CURRY  

SHLD CURRX  

SHLD CURRY :  

*****  

MNITR EQU $  

CALL NXCHR : GET COMMAND  

ADD A : OFFSET = 2*COMMAND  

MOV C, A :  

MVI B, 0 :  

LXI H, COMND : GET ADDRESS OF  

DAD B : COMMAND ROUTINE  

MOV E, M : INTO DE, THEN  

INX H :  

MOV D, M : INTO HL, THEN  

XCHG PCHL : INTO PC  

*****  

COMND EQU $  

DW PNLFT : MOVE W/ PEN UP  

DW PNDFP : MOVE W/ PEN DOWN  

DW XGRAF : X-AXIS GRAPH ROUTINE  

DW YGRAF : Y-AXIS GRAPH ROUTINE  

DW TXPLT : PLOT CHARACTERS  

DW RETCR : RETURN CURR. X,Y  

DW SETCR : RESET CURR. X,Y  

DW ENDIT : TERMINATES PLOTTING  

*****  

: COMMAND ROUTINES

```

PNLFT	EQU \$		
	CALL LIFT	;	LIFTS PEN IF NEC
	JMP PNMOV	;	#

PNDRP	EQU \$;	
	CALL DROP	;	DROPS PEN IF NEC
;			
PNMOV	EQU \$;	
	CALL REGIN	;	
	SHLD ENDX	;	
	CALL REGIN	;	
	SHLD ENDY	;	READS ENDPTS AND
	CALL PLOTS	;	MOVES PEN THERE
	JMP MNITR	;	

XGRAF	EQU \$;	PLOT X-AXIS GRAPH
	CALL NXCHR	;	GET X-INCR.
	MOV C, A	;	B IS CLEAR
	LHLD CURRX	;	MOVE PEN TO 1ST POINT
	SHLD ENDX	;	
	CALL REGIN	;	
	SHLD ENDY	;	
	CALL LIFT	;	
	CALL PLOTS	;	
LOOPX	EQU \$;	
	CALL REGIN	;	PLOT STREAM
	MOV A, H	;	GET NEXT Y POINT
	CPI NDCHR	;	CHECK FOR END OF STREAM
	JZ MNITR	;	
	SHLD ENDY	;	***
	LHLD CURRX	;	***
	DAD B	;	UPDATE X COORD
	SHLD ENDX	;	
	CALL DROP	;	***
	CALL PLOTS	;	INSURE PEN DOWN
	JMP LOOPX	;	***
	;	***	
	;	***	
YGRAF	EQU \$;	PLOT Y-AXIS GRAPH
	CALL NXCHR	;	GET Y-INCR.
	MOV C, A	;	B IS CLEAR
	LHLD CURRX	;	MOVE PEN TO 1ST POINT
	SHLD ENDX	;	
	CALL REGIN	;	
	SHLD ENDY	;	
	CALL LIFT	;	
	CALL PLOTS	;	
LOOPY	EQU \$;	
	CALL REGIN	;	PLOT STREAM
	MOV A, H	;	GET NEXT Y POINT
	CPI NDCHR	;	CHECK FOR END OF STREAM
	JZ MNITR	;	
	SHLD ENDY	;	***
	LHLD CURRX	;	***
	DAD B	;	UPDATE X COORD
	SHLD ENDX	;	
	CALL DROP	;	***
	CALL PLOTS	;	INSURE PEN DOWN
	JMP LOOPY	;	***
	;	***	
	;	***	
TXPLT	EQU \$;	
	CALL NXCHR	;	GET SCALE

```

MOV E, A           ; AND SAVE IN E
CALL NXCHR         ; GET DIRECTION
STA DIRM
***  

CLOOP EQU $          ; GET NEXT CHAR. CODE
CALL NXCHR         ; RETURN TO MONITOR
CPI NDCHR          ; AT END OF STRING
JZ MNITR
***  

LHLD CURRX         ; RESET ORIGINS
SHLD ORIGX         ;
LHLD CURRY         ;
SHLD ORIGY         ;
LXI H, CHTAB        ; DATA TABLE FOR CHAR GEN
ADD A              ; OFFSET=CODE+2
JNC ADOFF
***  

ADOFF INX H          ; A CONTAINS OFFSET
EQU $              ;
ADD L              ; #HL POINTS INTO CHTAB
MOV L, A            ; #
JNC GTSEG          ; #
INR H              ; #
GTSEG EQU $          ; BC POINTS TO SEG'S
MOV C, M            ;
INX H              ;
MOV B, M            ;
***  

SLOOP EQU $          ; #
LOAX B              ; GET NEXT SEGMENT CODE
MOV D, A              ; SAVE SEG. CODE IN D
CALL CHGEN          ; PLOT SEGMENT
INX B              ; INDEX OF NEXT SEGMENT
MOV A, D            ;
ANI EOCHR          ; #CHECK IF LAST SEGMENT
JZ SLOOP           ; CONT IF NOT LAST SEG
MVI A, NXORG        ; #PEN UP, (0,6) * SCALE
MOV D, A              ; INTERLETTER SPACE
CALL CHGEN          ; FOR NEXT CHAR
JMP CLOOP
***  

***  

RETCR EQU $          ; SEND CURRX,CURRY
LHLD CURRX
CALL PGOUT
LHLD CURRY
CALL RGOUT
***  

CALL SNDCR          ; SEND 'RET'
CALL HEXIN          ; WAIT FOR PROMPT
JMP MNITR
***  

SETCR EQU $          ; #
CALL REGIN
SHLD CURRX
CALL REGIN
SHLD CURPY
JMP MNITR
***  

ENDIT EQU R          ; RETURN TO MONITOR
RST 1
***  

: I/O UTILITIES
***  

REGIN EQU R          ; 4 HEX DIGITS INTO HL
CALL NXCHR
MOV H, A
CALL NXCHF
MOV L, A
RET
***
```

```

NXCHR EQU $ ; GETS HEX CONSTANT FROM BUFFER
PUSH B
PUSH D
PUSH H
LHLD PNTER ; CHECK IF BUFFER EMPTY
LXI D, BUFER
CALL CMPAF
DZ LOADR ; FILL IT IF NECESSARY
MOV B, M ; GET CHARACTER
INX H ;**
PUSH H
LXI D, BUFER
LHLD BUFLN ;**
DAD D
XCHG
POP H ;**
CALL CMPAR ;**
JN7 LDONE ;**
LDONE LXI H, BUFER : ELSE RESET PNTER
EQU $ ;**
SHLD PNTER
MOV A, B
POP H
POP D
POP B
RET

```

```

LOADR EQU   $          ; PREVENTS INK SPOTS
CALL LIFT
PUSH D
PUSH H
LXI H, BUFER
CALL SHAKE

CALL HEXIN      ; GET WORD COUNT
MOV D, A
CALL HEXIN
MOV E, A
XCHG :         ; SAVE BUFFER LENGTH
SHLD BUFLN
XCHG

LODNG EQU   $          ; GET NEXT WORD
CALL HEXIN
MOV M, A
INX H
DCX D
XRA A
ORA E
JNZ LODNG      ; CHECK IF LAST WORD
ORA D
JNZ LODNG
CALL HEXIN      ; REPEAT UNTIL DONE
                 ; WAIT FOR PROMPT
                 ; **

POP H
POP D
RET

```

```

MEXIN EQU $1000      ; 2 HEX DIGITS INTO H
PUSH B
CALL ASCHX          ; ASCII TO HEX CONVERSION
RLC
RLC
RLC
RLC
MOV B, A            ; SAVE HIGH CHARACTER
CALL ASCHX          ; GET LOW CHARACTER

```

```

        ORA    B          : THEN ADD
        POP    B
        RET

ASCHX EQU  $           : ASCII TO HEX IN A
CALL   RECEV          : GET ASCII CHAR
CPI    'A'             :
JC     ARNDA           : JUMP IF NUMBER
                      : ***

ARNDA ADI  9           : ADJUST A-F
EQU  $                 :
ANI   0FH              : CLEAR HIGH BITS
RET

                      : ***
                      : ***
                      : ***

RECEV EQU  $           : TEST MODEM STATS
IN    MDMCS            :
ANI   RXRDY            :
JZ    RECEV            : WAIT FOR CHAR
                      : ***
                      : READ CHAR
IN    MDMDT            :
ORA   A                :
JZ    RECEV            : IGNORE NULL CHARS
                      : ...LF
CPI   JAH              :
JZ    RECEV            :
CPI   GJH              : ...CR
JZ    RECEV            :
RET

                      : ***
                      : ***
                      : ***

RGOUT EQU  $           : (HL) TO 4 ASCII CHARS
CALL  HXOUT            :
MOV   H, L              :
CALL  HXOUT            :
RET

                      : ***
                      : ***
                      : ***

HXOUT EQU  $           : (H) TO 2 ASCII CHARS
MOV   A, H              :
RRC
RRC
RRC
RRC
CALL  HXASC            : LOW 4 BITS
MOV   A, H              :
CALL  HXASC            :
RET

                      : ***
                      : ***
                      : ***

HXASC EQU  $           : (A) TO 1 ASCII CHAR
ANI   0FH              : SEND LOW 4 BITS
CPI   0AH              :
JC    ARNDC            : CHECK FOR AND
                      : ***
                      : ADJUST A - F
ARNDC ADI  7           :
EQU  $                 :
ADI   "0"              : FOR ASCII "PREFIX"
CALL  XMIT             :
RET

                      : ***
                      : ***
                      : ***

XMIT  EQU  $           :
PUSH  B
MOV   C, A              :
                      : **

XLOOP EQU  $           :
IN    MDMCS            :
ANI   TXRDY            :
JZ    XLOOP            :

```

```

MOV A, C
OUT MDMDT
CALL RECEV
POP B
RET
***  

SHAKE EQU $  

PUSH H
MVI L, CAH      ; WAIT FOR CYBER FOX  

APPROX. 1 SEC
EWAIT EQU $  

CALL LONG
DCR L
JNZ EWAIT
POP H
MVI A, "1"      ; HANDSHAKE CHAR
CALL XMIT
SNDCR EQU $  

IN MDMCS      ; XMITS 'CR'  

ANI TXRDY
JZ SNDCR
MVI A, ODH      ; RETURN
OUT MDMDT      ; ECHO & LF IGNORED
***  

: PEN CONTROL UTILITIES
LIFT EQU $  

LDA PSTAT
ORA A
RZ
MVI A, LFTPN
CALL MOVE
JMP PWAIT
: CONTROLS PLOTTER
DROP EQU $  

LDA PSTAT
ORA A
RNZ
MVI A, DRPN
CALL MOVE
INR A          ; PSTAT GETS 1
: CONTROLS PLOTTER
PWAIT EQU $  

STA PSTAT
LONG EQU $  

PUSH H
MVI L, LWAIT
: L
LLOOP EQU $  

CALL SHORT
DCR L
JNZ LLOOP
POP H
RET
***  

MOVE EQU $  

OUT PLOTF      ; SEND INSTRUCTION
XRA A
OUT PLOTF      ; FOLLOWED BY 0
CALL SHORT     ; WAIT 1.65 MILLISEC

```

```

SHORT EQU $ :**
PUSH H
MVI L, SWAIT :**
WLOOP EQU $ :**
DCR L
JNZ WLOOP
POP H
RET :**
: ARITHMETIC UTILITIES :**
DPSUB EQU $ :#
MOV A, E :# DOUBLE PRECISION
SUB L :# SUBTRACTION ROUTINE
MOV L, A :#
MOV A, D :#
SBB H :#
MOV H, A :#
XRA A :# RESULT IN HL, ACCUM CLEARED
RET :#
:#
CMPAR EQU $ :#
MOV A, D :# DOUBLE PRECISION COMPARISON ROUTINE
CMP H :# HIGH BYTES NE 0
RNZ :#
MOV A, E :#
CMP L :#
RET :#
:#
SCALR EQU $ :#
ANI 7 :#
:#
MULT EQU $ :# 8 X 8 MULTIPLICATION
PUSH B :# ROUTINE
PUSH D :# HL GETS A+E
MVI C, 8 :#
MVI D, 0 :#
MOV L, D :#
MOV H, L :#
:#
MLOOP EQU $ :#
RAR :#
MOV B, A :#
JNC ARNDM :#
DAD D :#
CMC :#
:#
ARNDM EQU $ :#
XCHG H :# SHIFT MULTIPLICAND LEFT
DAD H :#
XCHG C :#
DCR C :#
JNZ MLOOP :#
:#
POP D :#
POP B :#
RET :#
: MAIN-LINE UTILITIES :**

```

```

CHGEN EQU $ :#
PUSH B :# SAVE REGS.
PUSH D :#
PUSH H :#
ANI 1000 :#
JNZ CHKD0N :# PEN UP/DOWN FIELD
              :# DECODE PEN UP/DOWN
CALL LIFT :#
JMP RELPT :#
:#
CHKDN EQU $ :#
CALL DROP :#

```

RELPT	EQU	\$		#
	MOV	A, D	D	CONTS SEG CODE
	CALL	SCALR	(A, AND 7)*SCALE	(E)
	MOV	B, H	LOADS BC WITH	
	MOV	C, L	RELATIVE Y	
:	MOV	A, D		#
	RRC			
	RRC			
	RRC			
	CALL	SCALR	SHIFT FOR RELATIVE X	
	XCHG		STORE IGNORES HIGH BITS IN A	
:	LDA	DIRN	COMPUTE DIRECTION OF TEXT	
	MOV	H, A	#	
	ORA	A	SETS FLAGS	
	JPE	DIRX	JUMP IF ON X-AXIS (0&3)	
:	PUSH	B		
	MOV	B, D	SWAP X AND Y VALUES	
	MOV	C, E	FOR VERTICAL LETTERS	
	POP	D	(1&2)	
**	DIRX	EQU	\$	#
	ANI	10B		#
	JZ	DIRY		#
:	XRA	A	COMP X VALUE	
	SUB	E	#	
	MOV	E, A	#	
	MVI	A, 0	#	
	SBB	D	#	
	MOV	D, A	DE GETS -DE	
**	DIRY	EQU	\$	#
	MOV	A, H		#
	ANI	01B		#
	JZ	SETPT		#
:	XRA	A	COMP Y VALUE	
	SUB	C	#	
	MOV	C, A	#	
	MVI	A, 0	#	
	SBB	B	#	
	MOV	B, A	BC GETS -BC	
**	SETPT	EQU	\$	#
:	LHLD	ORIGX		#
	DAD	D		#
	SHLD	ENDX		#
:	LHLD	ORIGY		#
:	DAD	B		#
:	SHLD	ENDY		#
:	CALL	PLOTS		#
:	POP	H	RESTORE REGS.	
:	POP	C		
:	POP	B		
:	RET			#
	CHTAB	EQU	R	DATA FOR CHAR GEN
	DW	CHR26	CHAR:	I
	DW	CHR33	CHAR:	A
	DW	CHR34	CHAR:	B
	DW	CHR35	CHAR:	C
	DW	CHR36	CHAR:	D
	DW	CHR37	CHAR:	E
	DW	CHR38	CHAR:	F
	DW	CHR39	CHAR:	G
	DW	CHR40	CHAR:	H
	DW	CHR41	CHAR:	I
	DW	CHR42	CHAR:	J

OW	CHR43	CHAR: X
OW	CHR44	CHAR: Z
OW	CHR45	CHAR: C
OW	CHR46	CHAR: V
OW	CHR47	CHAR: U
OW	CHR48	CHAR: W
OW	CHR49	CHAR: T
OW	CHR50	CHAR: J
OW	CHR51	CHAR: Y
OW	CHR52	CHAR: I
OW	CHR53	CHAR: O
OW	CHR54	CHAR: E
OW	CHR55	CHAR: S
OW	CHR56	CHAR: D
OW	CHR57	CHAR: F
OW	CHR58	CHAR: G
OW	CHR16	CHAR: H
OW	CHR17	CHAR: B
OW	CHR18	CHAR: N
OW	CHR19	CHAR: M
OW	CHR20	CHAR: P
OW	CHR21	CHAR: Q
OW	CHR22	CHAR: R
OW	CHR23	CHAR: A
OW	CHR24	CHAR: C
OW	CHR25	CHAR: L
OW	CHR11	CHAR: ;
OW	CHR13	CHAR: -
OW	CHR10	CHAR: *
OW	CHR15	CHAR: /SLASH
OW	CHR09	CHAR: (
OW	CHR04	CHAR:)
OW	CHR29	CHAR: =
OW	CHR69	CHAR: BLANK
OW	CHR12	CHAR: *
OW	CHR14	CHAR: #
OW	CHR03	CHAR: #POUND SIGN
OW	CHR59	CHAR: [OPEN SQ BRKT

OW	CHR61	CHAR:]
OW	CHR05	CHAR: PERCENT
OW	CHR02	CHAR: "DBL. QUOTE
OW	CHR63	CHAR: UNDERSCORE
OW	CHR01	CHAR: EXCLAMATION
OK	CHR06	CHAR: AMPERSAND
OW	CHR07	CHAR: •APOSTROPHE
OW	CHR31	CHAR: ?QUESTION MARK
OW	CHR28	CHAR: <
OW	CHR30	CHAR: >
OW	CHR32	CHAR: @COMMERCIAL AT
OW	CHR60	CHAR: /\BACK SLASH
OW	CHR62	CHAR: ^CIRCUMFLEX
OW	CHR27	CHAR: ..
OW	CHR54	CHAR: APOSTROPHE
OW	CHR65	CHAR: OPEN PAREN
OW	CHR66	CHAR: VERTICAL LINE
OW	CHR67	CHAR: CLOSE PAREN
OW	CHR68	CHAR: TILDE

CHLST EQU \$: CHARACTER TABLES

CHR01 EQU \$: CHAR:

DB	0160
DB	1120
DB	0110
DB	3100

:**

CHR02 EQU \$: CHAR: "

DB	0140
DB	1260
DB	0340
DB	3460

:**

CHR03 EQU \$: CHAR: #

DB	0130
----	------

DB 1260
DB 0360
DB 1230
DB 0350
DB 1150
DB 0140
DB 3340

;**

CHR04 EQU \$; CHAR: \$

DB 0020
DB 1110
DB 1310
DB 1420
DB 1330
DB 1130
DB 1040
DB 1150
DB 1350
DB 1460
DB 0260
DB 3200

;**

CHR05 EQU \$; CHAR: PERCENT

DB 1460
DB 1160

DB 1750
DB 1140
DB 1250
DB 1160
DB 0410
DB 1320
DB 1210
DB 1300
DB 3410

;**

CHR06 EQU \$; CHAR: &

DB 0420
DB 1200
DB 1100
DB 1010
DB 1320
DB 1350
DB 1260
DB 1150
DB 3400

;**

CHR07 EQU \$; CHAR: ''

DB 0260
DB 3240

;**

CHR08 EQU \$; CHAR: (

DB 0360
DB 1240
DB 1220
DB 3300

;**

CHR09 EQU \$; CHAR:)

DB 0160
DB 1240
DB 1220
DB 3100

;**

CHR10 EQU \$; CHAR: *

DB 0130
DB 1350
DB 0250
DB 1230
DB 0150
DB 3330

;**

```

:: CHR11 EQU $ : CHAR: +
DB 0130
DB 1330
DB 0240
DB 3220
::+
:: CHR12 EQU $ : CHAR: ,
DB 9220
DB 3100
::+
:: CHR13 EQU $ : CHAR: -
DB 0130
DB 3330
::+
:: CHR14 EQU $ : CHAR: .
DB 0110
DB 3100
::+
:: CHR15 EQU $ : CHAR: :
DB 3460
::+
:: CHR16 EQU $ : CHAR: 0
DB 0100
DB 1020
DB 1940
DB 1160
DB 1360
DB 1440
DB 1420
DB 1300
DB 3100
::+
:: CHR17 EQU $ : CHAR: 1
DB 0140
DB 1260
DB 3200
::+
:: CHR18 EQU $ : CHAR: 2
DB 0050
DB 1160
DB 1360
DB 1450
DB 1440
DB 1120
DB 1000
DB 3400
::+
:: CHR19 EQU $ : CHAR: 3
DB 0010
DB 1100
DB 1330
DB 1410
DB 1420
DB 1330
DB 1230
DB 0330
DB 1440
DB 1450
DB 1360
DB 1160
DB 3050
::+
:: CHR20 EQU $ : CHAR: 4
DB 0300
DB 1360
DB 1020
DB 3420

```

CHR21 EQU \$; CHAR: 5

DB 0010
DB 1100
DB 1300
DB 1410
DB 1430
DB 1340
DB 1140
DB 1030
DB 1060
DB 3460

;**

CHR22 EQU \$; CHAR: 6

DB 0030
DB 1140
DB 1340
DB 1430
DB 1410
DB 1300
DB 1100
DB 1010
DB 1050
DB 1160
DB 1360
DB 3450

;**

CHR23 EQU \$; CHAR: 7

DB 0060
DB 1460
DB 3100

;**

CHR24 EQU \$; CHAR: 8

DB 0100
DB 1010
DB 1020
DB 1130
DB 1040
DB 1050
DB 1160
DB 1360
DB 1450
DB 1440
DB 1330
DB 1130
DB 0330
DB 1420
DB 1410
DB 1300
DB 3100

;**

CHR25 EQU \$; CHAR: 9

DB 0010
DB 1100
DB 1300
DB 1410
DB 1450
DB 1360
DB 1160
DB 1050
DB 1030
DB 1120
DB 1320

;**

DB 3430

;**

CHR26 EQU \$; CHAR: 8

DB 0140
DB 1130

```

    DB 0120
    DB 3110      ;**
    CHR27 EQU $      : CHAR: :
    DB 0240
    DB 1230
    DB 0220
    DB 3100      ;**
    CHR28 EQU $      : CHAR: <
    DB 0460
    DB 1030
    DB 3400      ;**
    CHR29 EQU $      : CHAR: =
    DB 0040
    DB 1440
    DB 3020
    DB 3420      ;**
    CHR30 EQU $      : CHAR: >
    DB 1430
    DB 3060      ;**
    CHR31 EQU $      : CHAR: ?
    DB 0050
    DB 1160
    DB 1260
    DB 1350
    DB 1340
    DB 1130
    DB 1120
    DB 0110
    DB 3100      ;**
    CHR32 EQU $      : CHAR: @
    DB 0330
    DB 1220
    DB 1130
    DB 1240
    DB 1340
    DB 1320
    DB 1430
    DB 1440
    DB 1350
    DB 1150
    DB 1040
    DB 1020
    DB 1110
    DB 1310
    DB 3420      ;**
    END1 EQU $      ;**  

    ;  

    ;  

    ;  

    ;  

    ORG 0000H      ;**  

    ;  

    ;  

    PLOTS EQU $      : #  

    PUSH 3  

    PUSH 0  

    PUSH H      : #  

    ;  

    LXI B, C      : CONVIENT CONSTANTS
    ;  

    LHLD CURPX      : #
    XCHG  

    LHLD ENOX      : LOADS DE WITH CURPX AND RESETS CURR
    SHLD CUPPX      : #

```

```

XCHG
;
CALL CMPAR      ; COMPARES (DE) TO (HL)
JC NEGDX
JN7 POSDX      ; ENDX<CURRX
;
NODX: CALL YCOMP ; ENDX>CURRX
JC QUAD4
JNZ QUAD2
JMP WRAP       ; ENDY<CURRY
;
POSDX EQU $      ; #
CALL YCOMP
JC QUAD4
JMP QUAD1       ; ENDY<CURRY
;
NEGDX EQU $      ; #
XCHG           ; #
CALL YCOMP
JC QUAD3
JNZ QUAD2
JMP QUAD3       ; ENDY>=CURRY
;
QUAD4: RLC      ; #
INR B
QUAD3: RLC      ; #
INR B
XCHG           ; KEEPS DX POSITIVE
QUAD2: RLC      ; #(DE)=GTR(ENDY, CURRY)
INR B            ; #(HL)=LSR(CURRY, ENDY)
QUAD1: RLC      ; #
INR B
ANI 0FH
STA DIAG
CALL OPSUR
SHLD DY         ; USE LOWEST 4 BITS
; COMPUTES DY=(DE)-(HL)
; COMPUTES DY (DE)-(HL)
; AND CLEARS ACCUM.

;
MOV D, C        ; #(DE)=0
MOV E, C
XCHG           ; #
SHLD TEST
SHLD TEST + 2
SHLD COMP
SHLD COMP + 2   ; CLEARS VARIABLES
STA MAJ
CALL CMPAR     ; COMPARES (DE)&(HL)... 0 & DY
;
```

```

JZ SEC5          ; DY = ZERO
XCHG           ; #
LHLD DX          ; #
XCHG           ; #
CALL CMPAR
JZ SEC5          ; DX=ZERO
;
LHLD DY          ; #
CALL CMPAR
JC HISEC         ; DX<DY, M>1
;
MVI A, 1          ; DX>=DY, M<=1
STA MAJ
;
LDA DIAG
ANI 0101B
STA HORV
DAD H
CALL CMPAR
JC SEC2
JMP SEC1         ; DX<2*DY, .5<M<1
; DX>=2*DY, 0<M<.5
;
HISEC EQU $      ; #
LDA DIAG
ANI 1010B
STA HORV
XCHG           ; COMPUTES HORV
DAD H
CALL CMPAR
JC SEC3          ; #(DE)=DY, (HL)=DX
; DX>2*DX, 1<M<2
;
```

```

        JMP SEC4      ; DY>=2*DX, M>2
: SEC5 EQU $          ; #
XRA A          ; A IS 0
STC             ; **

LOOP5 EQU $          ; #
RAL             ; #
DCR B          ; LOOP COMPUTES
JNZ LOOPS      ; # PLOT CHAR
STA HORV       ; # FROM QUADRANT (B)
MVI C, 4        ;
JMP SDRIV      ;

: SEC4: INR C          ; #
SEC3: INR C          ; #
SEC2: INR C          ; #
SEC1: LDA DIAG       ; PROCESS QUADS 3, 4 SEPARATELY
ANI 1000B       ; #
JNZ QAD34       ; #
MOV A, C          ; #
JMP SDRIV      ; #

: QAD34: MVI A, 3      ; #
SUB C           ; #

SDRIV EQU $          ; #
STA DRIVR      ; #

: LINSET          ; #
LHLD DY          ; #
XCHG            ; #
LHLD DX          ; # (DE)=DY, (HL)=DX
LOA MAJ         ; #
DCR A           ; #

: XMAJ JNZ XMAJ      ; PROCESS ABS(SLOPE)>1 SEPARATELY
XCHG            ; #
EQU $           ; #
SHLD OXI         ; #
SHLD IDIFF      ; #
MOV A, C          ; #
MOV B, H          ; #
MOV C, L          ; #
DAD H           ; #
XCHG            ; # (DE)=2*DXI, (HL)=DXJ
SHLD DXJ         ; #
SHLD JOIFF      ; #

: SUI 4           ; A=OLD C (SECTOR)
JZ XSDRV        ; TRAP OUT AXES
PUSH H           ; #(BC)=IDIFF, ((SP))=JDIFF
LOA DRIVR       ; #
ANI 1           ; #
JNZ LNTST       ; PROCESS DRIVERS 1, 3 SEPARATELY
SHLD COMP        ; #
XCHG            ; #
SHLD TEST        ; # (DE)=DXJ, (HL)=2*DXI
TEST=2*IDIFF    ; # TEST=2*IDIFF

: LNTST EQU $          ; #
XRA A           ; CLEAR A
LXI H, OXI       ; #
ORA M           ; #
INX M           ; #
ORA M           ; #
INX H           ; #
ORA M           ; #
INX H           ; #
ORA M           ; #
JNZ LNCMP        ; IF((DXI.OR.DXJ).EQ.0)
POP D           ; CLFARS STACK
JMP WRAP        ; RETURNS TO MONITOR
: **

LNCMP EQU $          ; #

```

```

;                                         (TEST-COMP)
LHLD TEST + 2
XCHG
LHLD COMP + 2
CALL CMPAR
JC PHORV
JNZ PDIAG

LHLD TEST
XCHG
LHLD COMP
CALL CMPAR
JC PHORV
JZ EQTST

PDIAG EQU $          ;#
LDA  DIAG
CALL MOVE
LHLD DXI
DCX  H
SHLD DXI
LHLD COMP
POP  D          ;# SEND TO PLOTTER
DXI=DXI-1

PUSH D          ;# BUT LEAVE ON STACK
DAD D
SHLD COMP
JNC PWRAP

LHLD COMP + 2
INX  H
SHLD COMP + 2
JMP  PWRAP
EQTST EQU $          ;#
LDA  DRIVR
DCR  A
JZ   PDIAG
POP  D
PUSH D          ;# BUT LEAVE ON STACK
MOV  H, B
MOV  L, C
DAD B
CALL CMPAR
LDA  DRIVR
JNZ  TSTD1
CPI  3
JZ   PDIAG
JMP  PHORV
TSTD1 EQU $          ;#
DCR  A
JM   PDIAG

PHORV EQU $          ;#
LDA  HORV
CALL MOVE
PWRAP EQU $          ;# SEND TO PLOTTER
LHLD DXJ
DCX  H
SHLD DXJ
LHLD TEST
DAD  9
SHLD TEST
JNC  LNTST
LHLD TEST + 2
INX  H
SHLD TEST + 2
JMP  LNTST

XSDRV EQU $          ;#
DAD  B
LDA  HORV
MOV  C, A          ;# ADD (BC)=IDIFF TO (HL)=JOIFF
LOAD C WITH HORV
PAXIS EQU $          ;# SEND TO PLOTTER
MOV  A, C
CALL MOVE
DCX  H
ORA  L          ;# A WAS CLEAR

```

```

JNZ  PAXIS
ORA H
JNZ  PAXIS
      ;**+
      ;**+
      ;**+
WRAP EQU $          ; RESTORE
POP H
POP D
POP B
      ;**+
      ;**+
      ;**+
RET
      ;**+
YCOMP EQU $          ; # COMPUTE DX
CALL OPSUB
SHLD DX
      ;**+
LHLD CURRY
XCHG
LHLD ENDY
SHLD CURRY
XCHG
CALL CMPAF
MVI A, DICMP
RET      ;# LOADS A WITH 10010001B
      ;#;
      ;#;
CHR33 EQU $          ; CHAR: A
DB 1050
DB 1160
DB 1360
DB 1450
DB 1400
DB 0030
DB 3430
      ;**+
      ;**+
CHR34 EQU $          ; CHAR: B
DB 1060
DB 1360
DB 1450
DB 1440
DB 1330
DB 1030
DB 0330
DB 1420
DB 1410
DB 1300
DB 3000
      ;**+
      ;**+
CHR35 EQU $          ; CHAR: C
DB 0450
DB 1360
DB 1160
DB 1050
DB 1010
DB 1100
DB 1300
DB 3410
      ;**+
      ;**+
CHR36 EQU $          ; CHAR: D
DB 1060
DB 1360
DB 1450
DB 1410
DB 1300
DB 3000
      ;**+
      ;**+
CHR37 EQU $          ; CHAR: E
DB 1060
DB 1460
DB 0230
DB 1030

```

	DB	0000	
	DB	3400	
			:**
CHR38	EQU	\$: CHAR: F
	DB	1060	
	DB	1460	
	DB	0230	
	DB	3030	
			:**
CHR39	EQU	\$: CHAR: G
	DB	0230	
	DB	1430	
	DB	1410	
	DB	1300	
	DB	1100	
	DB	1010	
	DB	1050	
	DB	1160	
	DB	1360	
	DB	3450	
			:**
CHR40	EQU	\$: CHAR: H
	DB	1050	
	DB	0030	
	DB	1430	
	DB	0460	
	DB	3400	
			:**
CHR41	EQU	\$: CHAR: I
	DB	1400	
	DB	0460	
	DB	1060	
	DB	0260	
	DB	3200	
			:**
CHR42	EQU	\$: CHAR: J
	DB	0020	
	DB	1010	
	DB	1100	
	DB	1300	
	DB	1410	
	DB	3460	
			:**
CHR43	EQU	\$: CHAR: K
	DB	1060	
	DB	0020	
	DB	1460	
	DB	0130	
	DB	3400	
			:**
CHR44	EQU	\$: CHAR: L
	DB	0060	
	DB	1000	
	DB	3400	
			:**
CHR45	EQU	\$: CHAR: M
	DB	1060	
			:**
	DB	1230	
	DB	1460	
	DB	3400	
			:**
CHR46	EQU	\$: CHAR: N
	DB	1060	
	DB	1400	
	DB	3460	
			:**

```

CHR47 EQU $ ; CHAR: O
DB 0010
DB 1050
DB 1160
DB 1360
DB 1450
DB 1410
DB 1300
DB 1100
DB 3010
;**
CHR48 EQU $ ; CHAR: P
DB 1060
DB 1360
DB 1450
DB 1440
DB 1330
DB 3030
;**
CHR49 EQU $ ; CHAR: Q
DB 0010
DB 1050
DB 1160
DB 1360
DB 1450
DB 1410
DB 1300
DB 1100
DB 1010
DB 0310
DB 3400
;**
CHR50 EQU $ ; CHAR: R
DB 1060
DB 1360
DB 1450
DB 1440
DB 1330
DB 1030
DB 0130
DB 3400
;**
CHR51 EQU $ ; CHAR: S
DB 0010
DB 1100
DB 1300
DB 1410
DB 1420
DB 1330
DB 1130
DB 1040
DB 1050
DB 1160
DB 1360
DB 3450
;**
CHR52 EQU $ ; CHAR: T
DB 0060
DB 1460
DB 0260
DB 3200
;**
;**
CHR53 EQU $ ; CHAR: U
DB 0060
DB 1010
DB 1100
DB 1300
DB 1410

```

	DB	3460	;
CHR54	EQU	\$;
	DB	0060	CHAR: V
	DB	1280	
	DB	3460	
			;
CHR55	EQU	\$;
	DB	0060	CHAR: W
	DB	1060	
	DB	1230	
	DB	1400	
	DB	3460	
			;
CHR56	EQU	\$;
	DB	0060	CHAR: X
	DB	1460	
	DB	0060	
	DB	3400	
			;
CHR57	EQU	\$;
	DB	0060	CHAR: Y
	DB	1230	
	DB	1460	
	DB	0230	
	DB	3280	
			;
CHR58	EQU	\$;
	DB	0060	CHAR: Z
	DB	1460	
	DB	1000	
	DB	3400	
			;
CHR59	EQU	\$;
	DB	0460	CHAR: [
	DB	1260	
	DB	1200	
	DB	3400	
			;
CHR60	EQU	\$;
	DB	0060	CHAR: \
	DB	3400	
			;
CHR61	EQU	\$;
	DB	1200	CHAR:]
	DB	1260	
	DB	3060	
			;
CHR62	EQU	\$;
	DB	0030	CHAR: ^
	DB	1260	
	DB	3430	
			;
CHR63	EQU	\$;
	DB	3400	CHAR: _
			;
CHR64	EQU	\$;
	DB	0160	CHAR:
	DB	3240	
			;
CHR65	EQU	\$;
	DB	0460	CHAR:
	DB	1350	
	DB	1340	
	DB	1230	
	DB	1320	

```

DB    1310
DB    3400
;**
CHR66 EQU $ ; CHAR:
DB    0250
DB    3210
;**
CHR67 EQU $ ; CHAR:
DB    0100
DB    1210
DB    1220
DB    1330
DB    1240
DB    1250
DB    3160
;**
CHR68 EQU $ ; CHAR:
DB    0020
DB    1130
DB    1310
DB    3420
;**
CHR69 EQU $ ; CHAR: BLANK
DB    2000
;**
;**
END2 EQU $-1

```

```

        ORG 1000H
ENDX: DW 0
ENDY: DW 0
CURRX: DW 0
CURRY: DW 0
ORIGX: DW 0
ORIGY: DW 0
DX: DW 0
DY: DW 0
DXI: DW 0
DXJ: DW 0
IDIFF: DW 0
JDIFF: DW 0
;
DIAG: DB 0
HORV: DB 0
DRIVR: DB 0
MAJ: DB 0
SCALE: DB 0
DIRN: DB 0
PSTAT: DB 0
DSTAT EQU 1
USTAT EQU 0
;;
TEST: DW 0
COMP: DW 0
DW 0
PNTER: DW 0
BUFLN: DW 0
;
BUFER EQU $1000
LSTND EQU S-1
END
;
```

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